

Lumière sur le côté obscur de l'Univers



Peter von Ballmoos, IRAP Toulouse

la lumière qui nous parvient des astres
nous inspire bon nombre "d'idées noires" ...

ciel nocturne noir

nébuleuses sombres

trous noirs

matière noire

énergie sombre

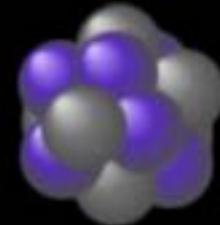
INTERNATIONAL
YEAR OF
LIGHT
2015



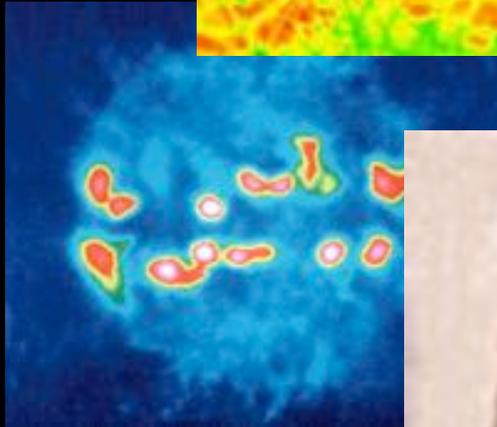
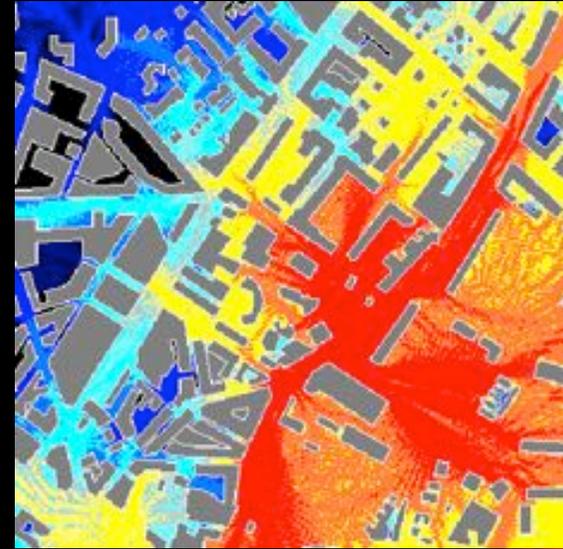
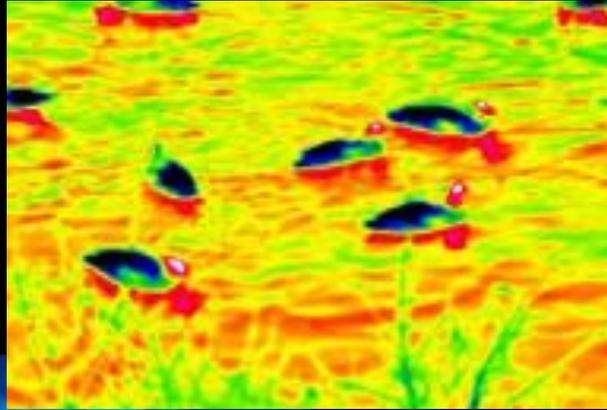


lumière

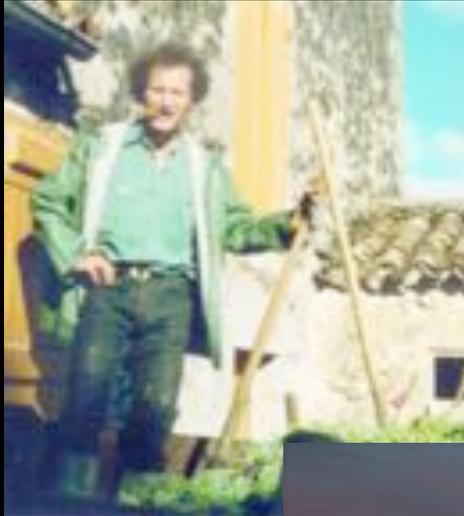
visible et invisible



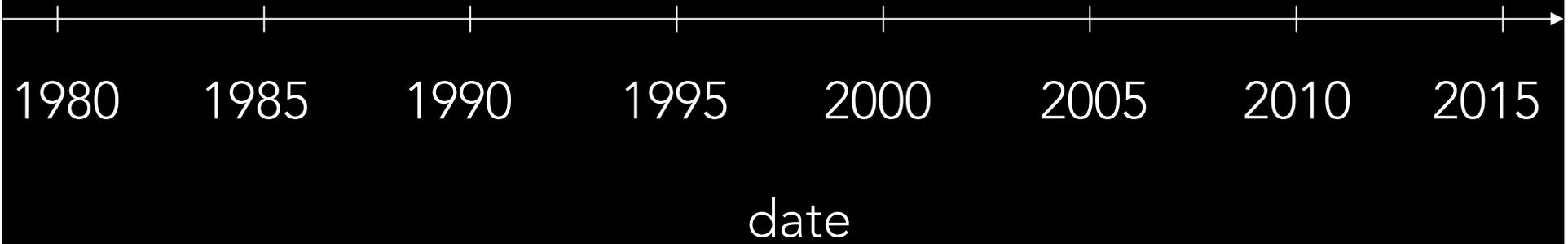
Imaginer l'Univers invisible ?



Comment imaginer *mon* Univers ?



Comment classer les images de *mon* Univers ?



date

Comment classer les images de *mon* Univers ?



cm

m

km

taille

Comment classer les images de *mon* Univers ?



rouge

orange

jaune

vert

bleu

violet

couleurs

lumière visible ?



lumière visible ?

*Slide your hand inside.
Visible light is can only be seen
when an object is present.
Desliza tu mano al interior.
La luz visible sólo se percibe por
la presencia de un objeto.*

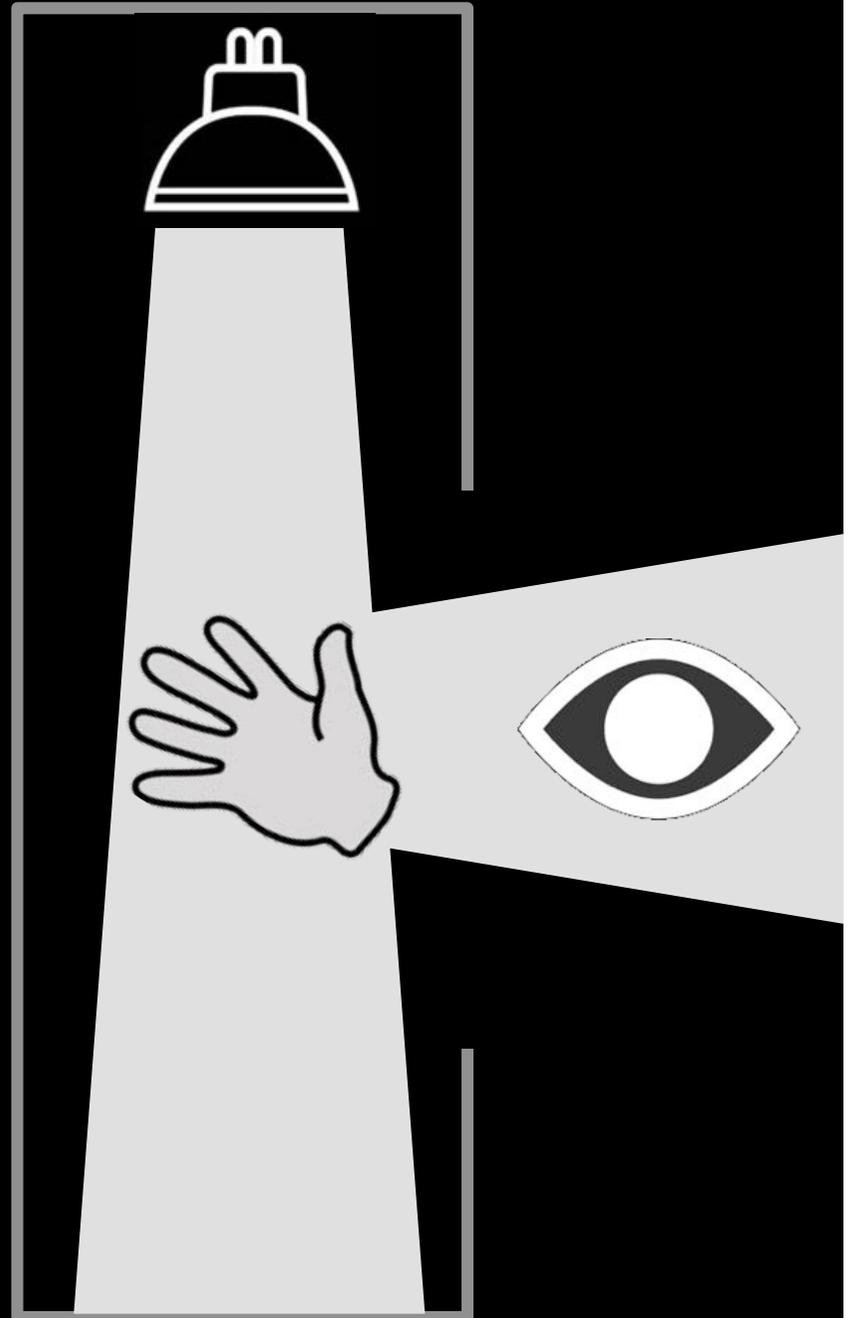
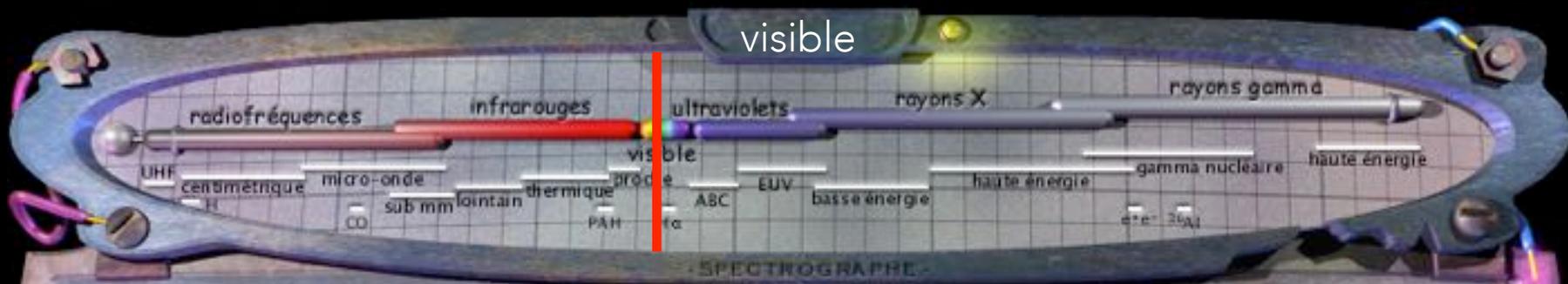
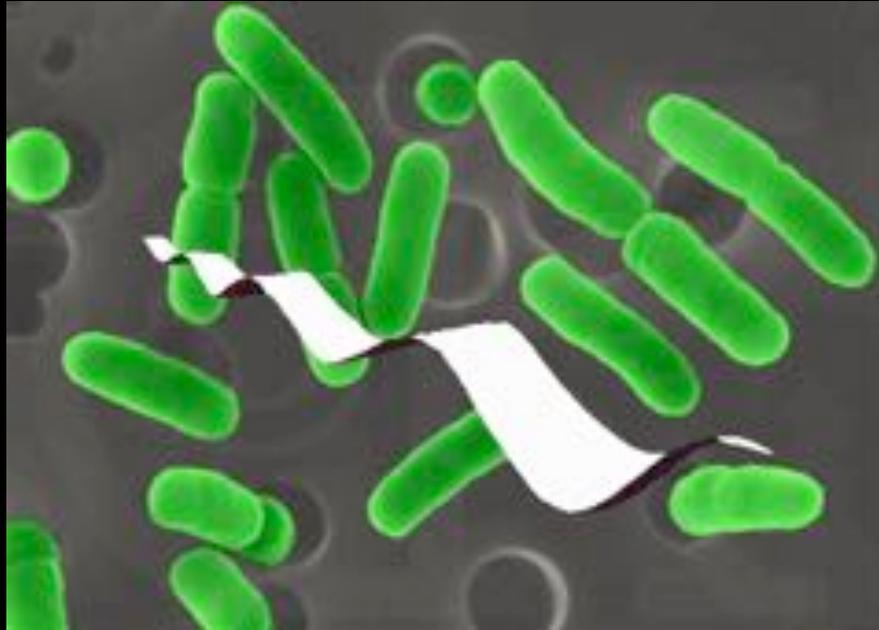
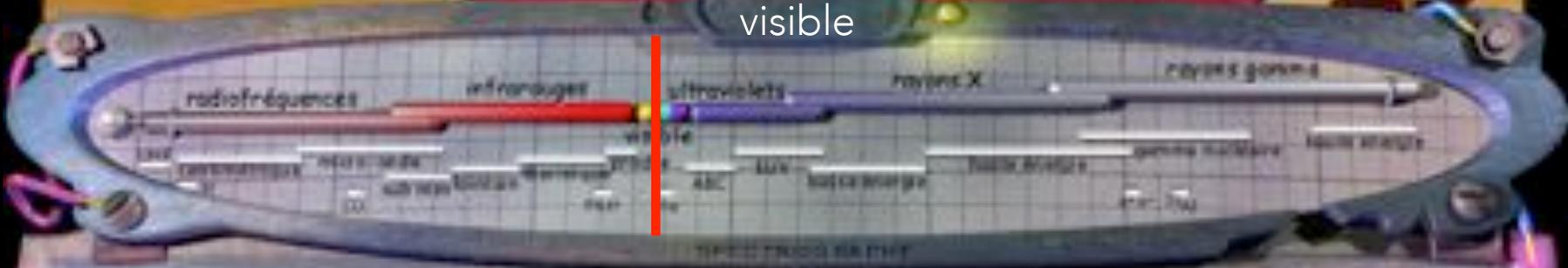


tableau de bord : lumière visible



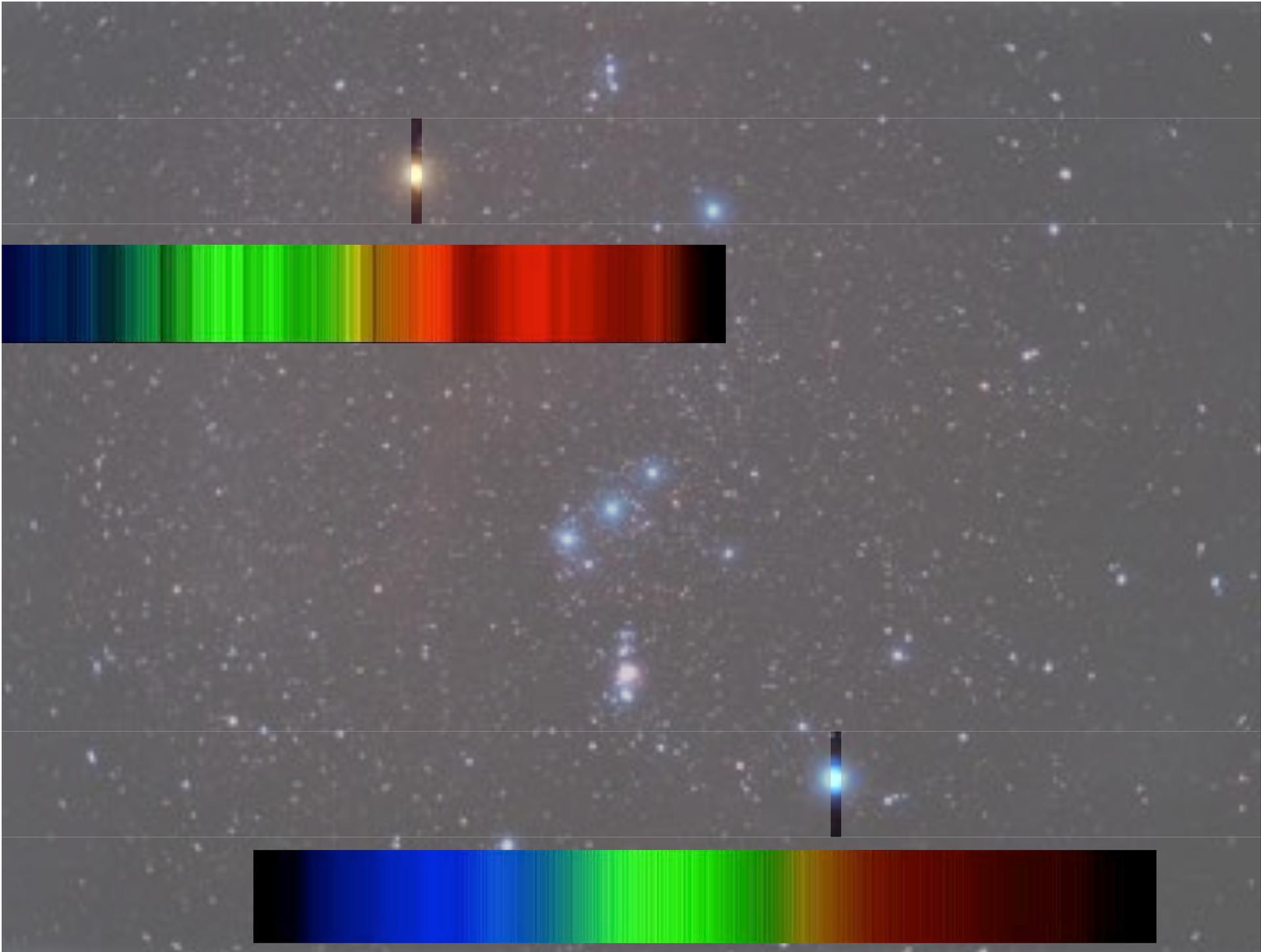


visible

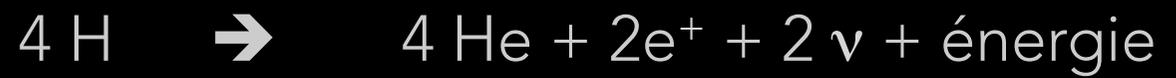
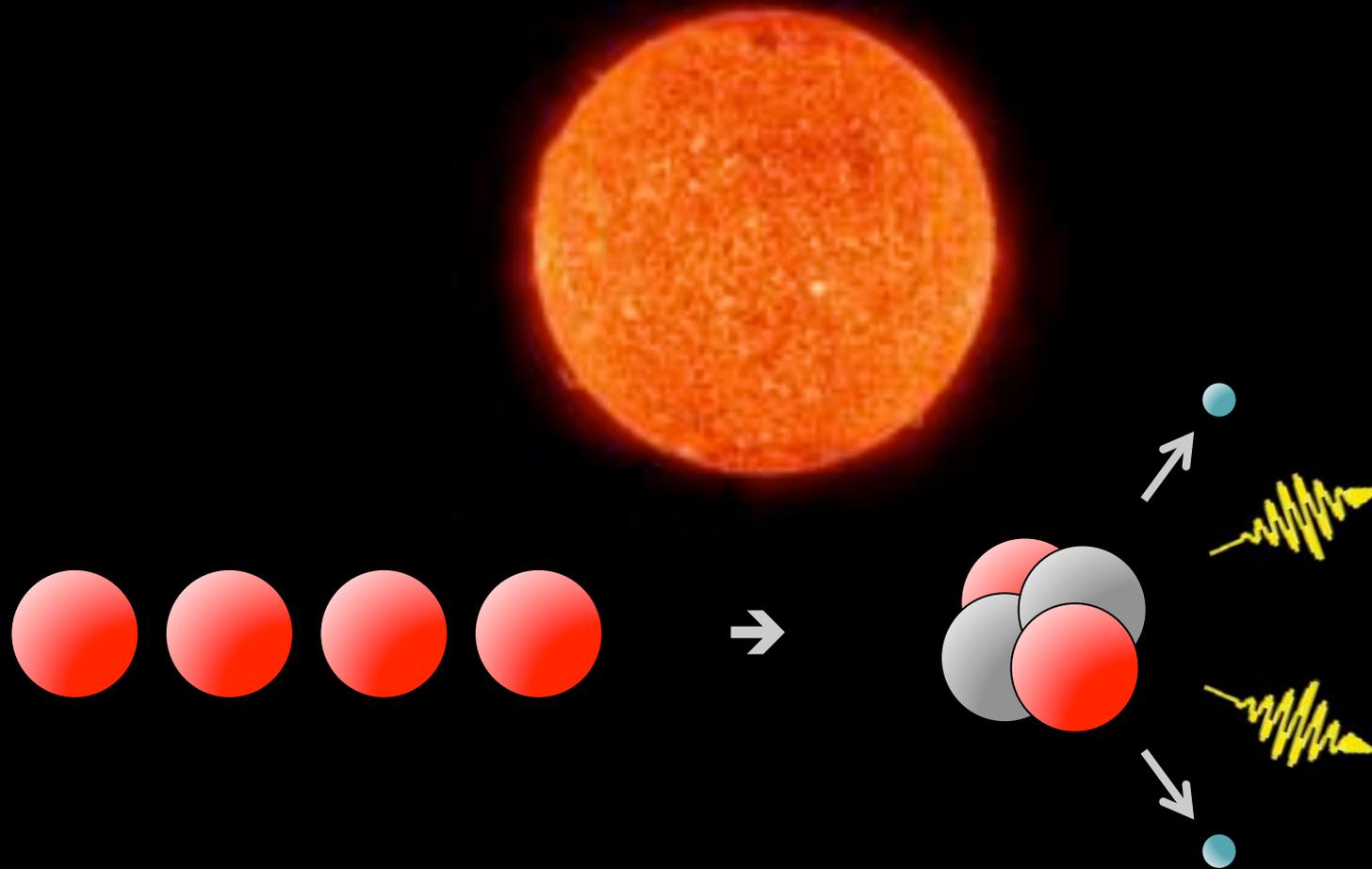




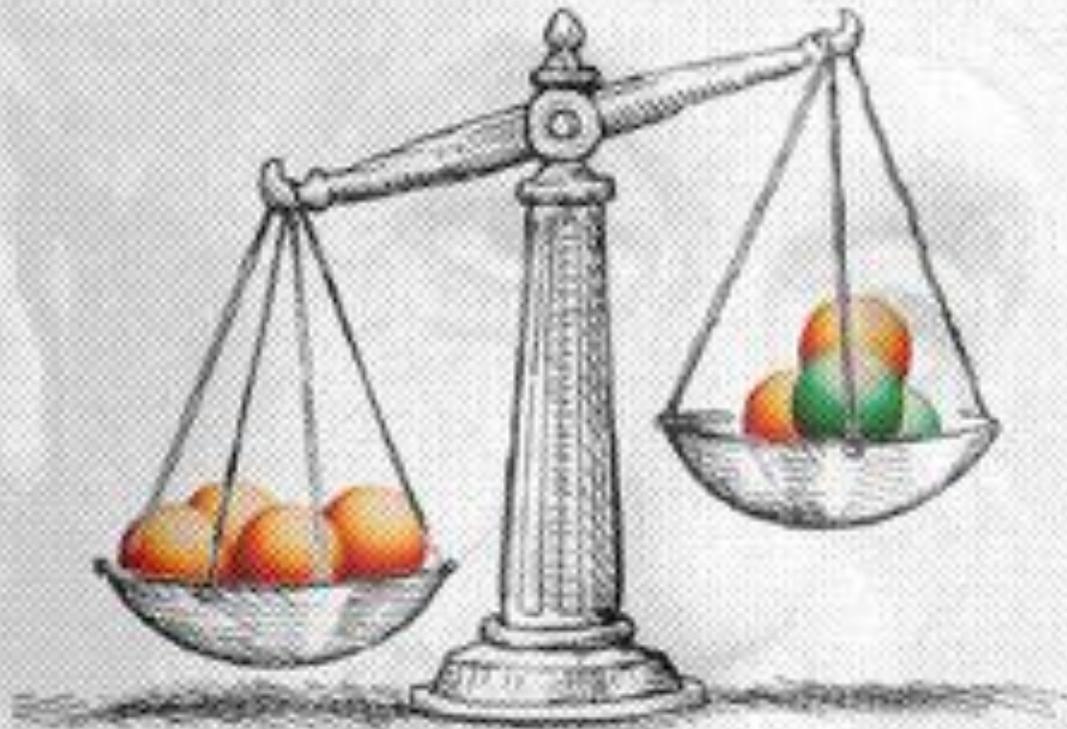




La fusion H \rightarrow He dans le Soleil



$$E = mc^2$$

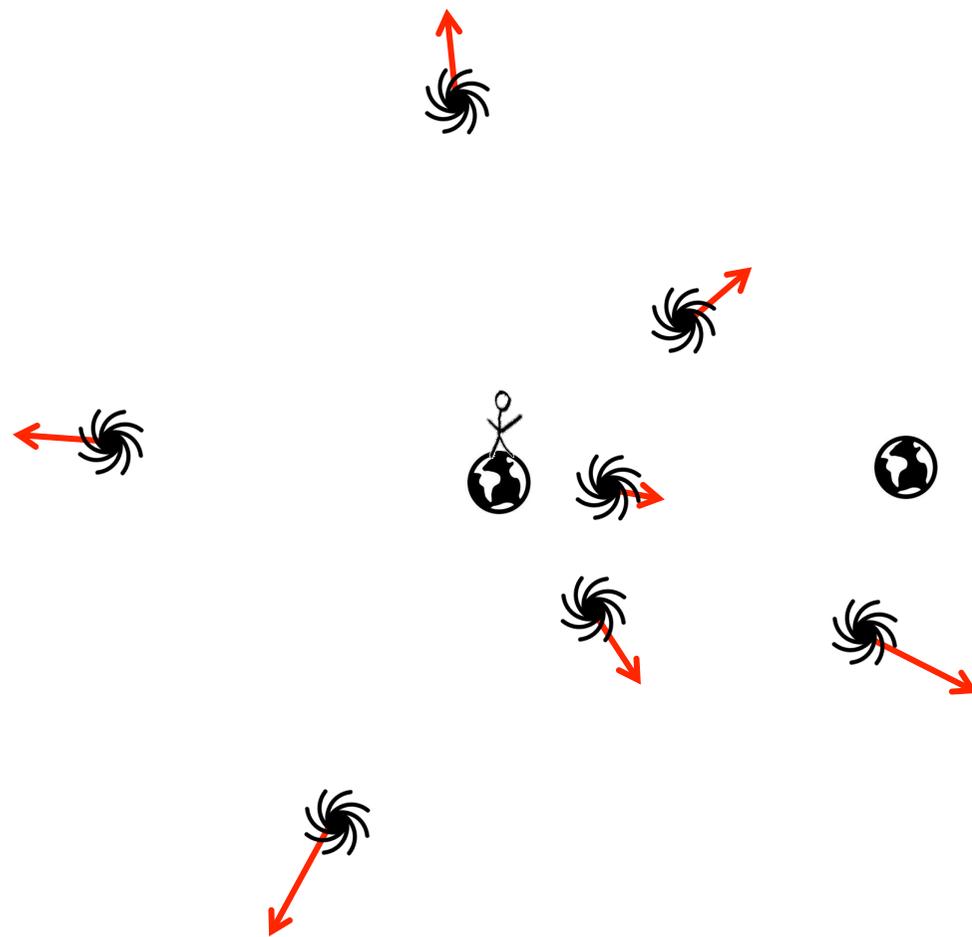


} Δm

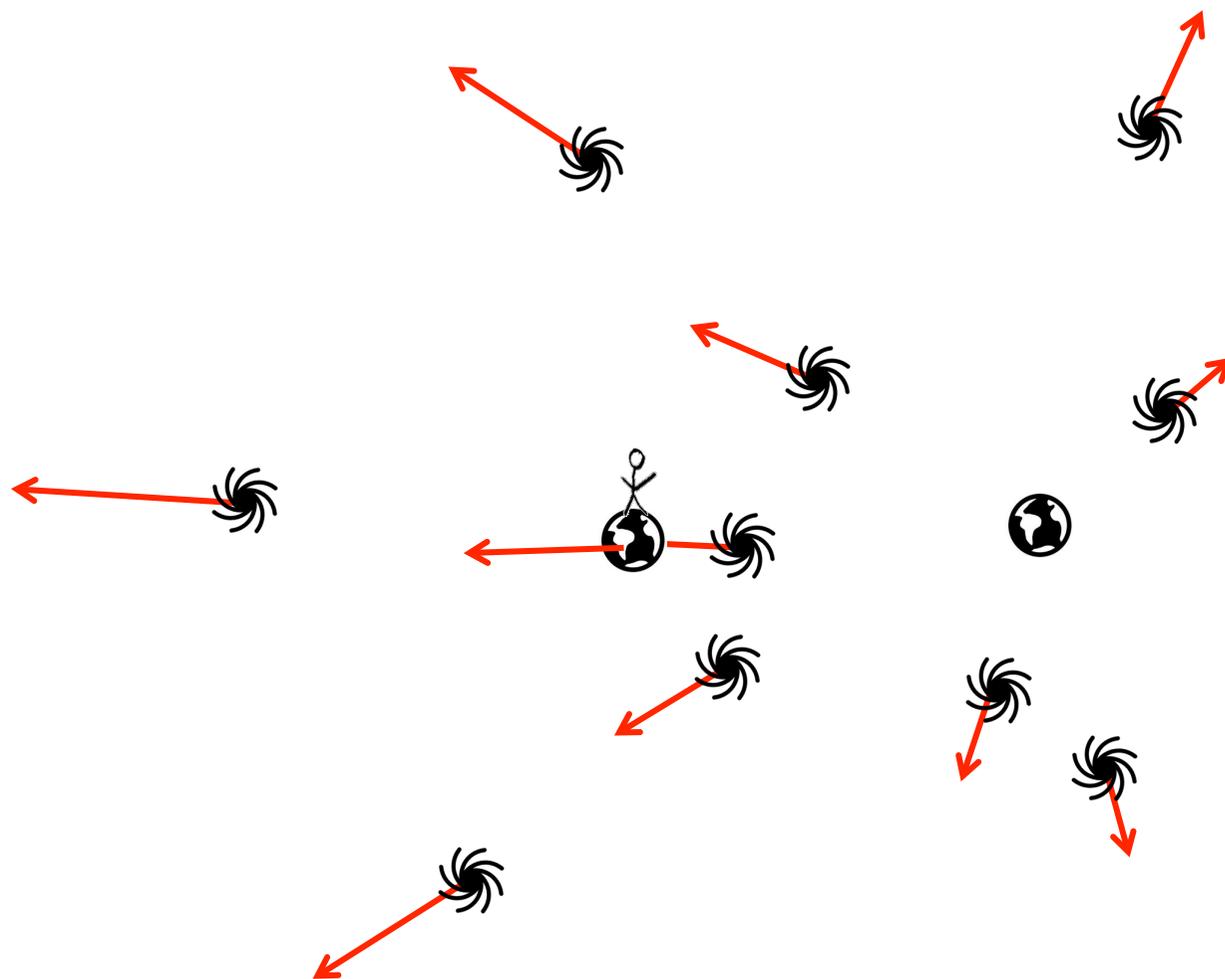


la récession des galaxies

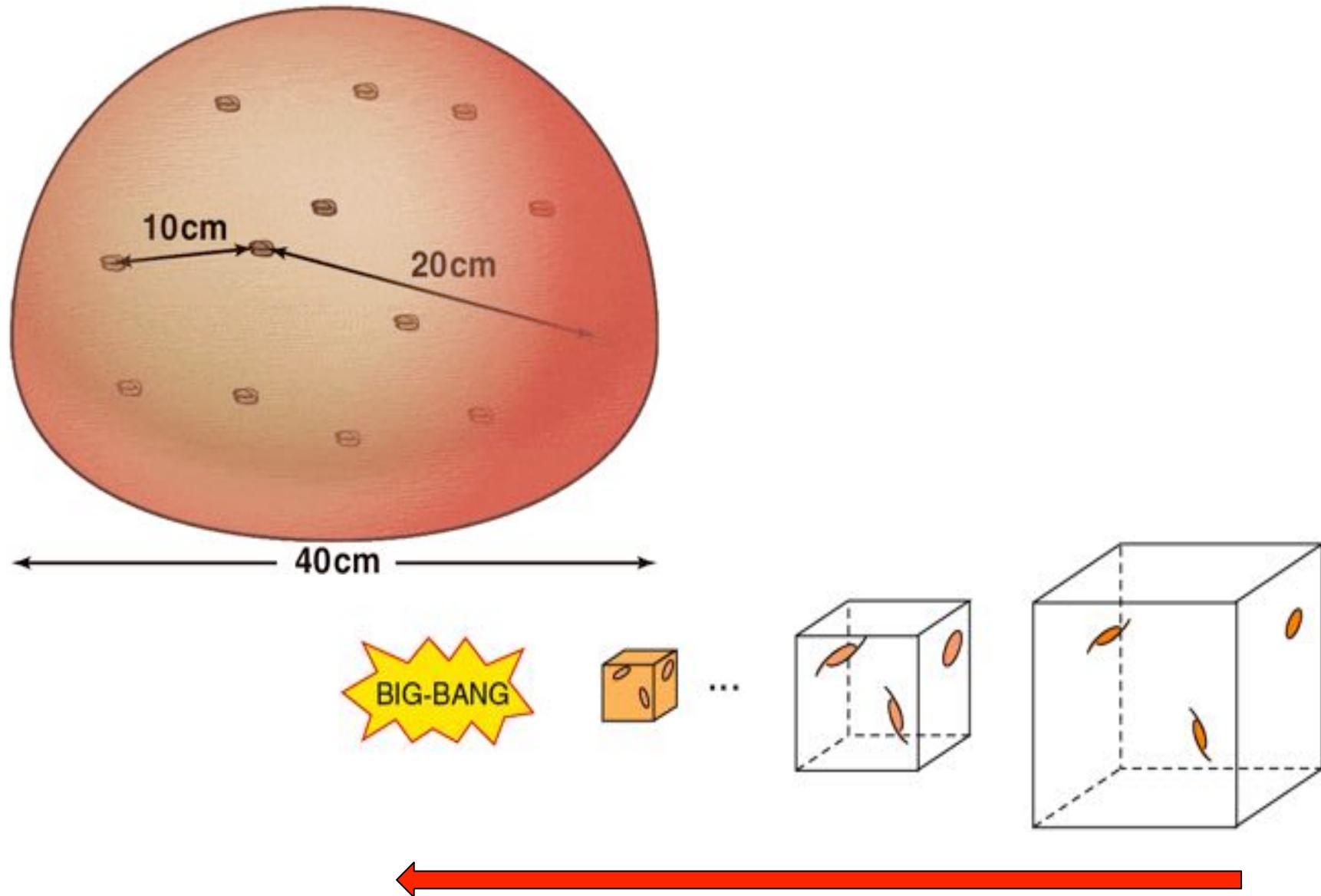
la récession des galaxies



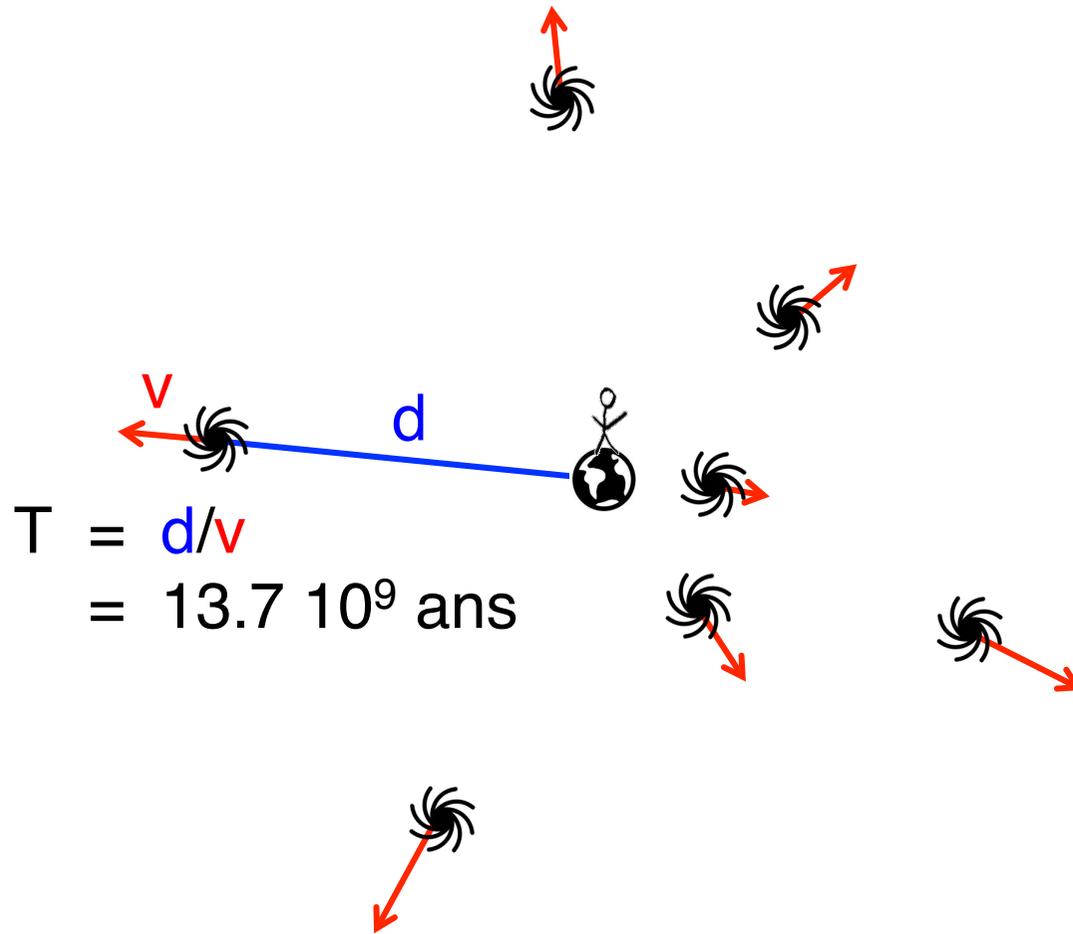
la récession des galaxies



l'expansion de l'Univers



l'expansion de l'Univers



lumière visible ...



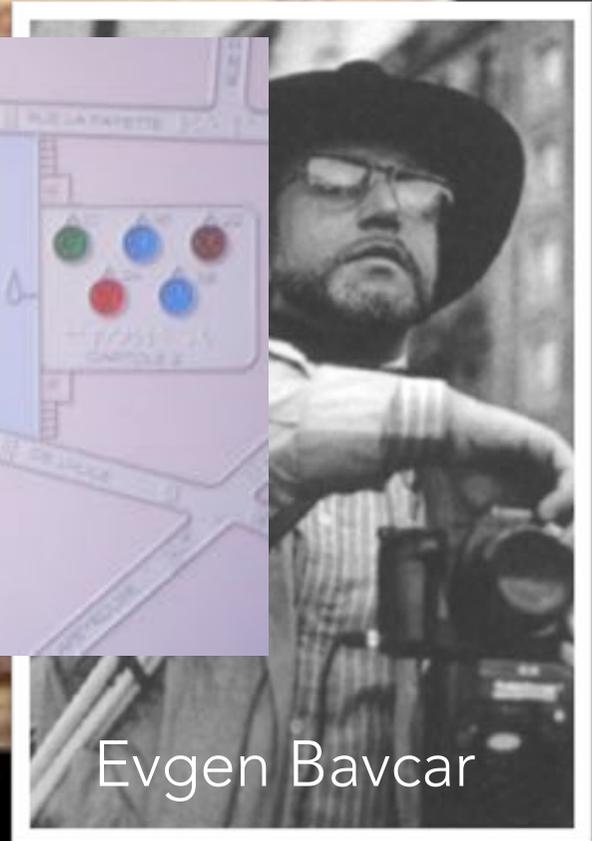
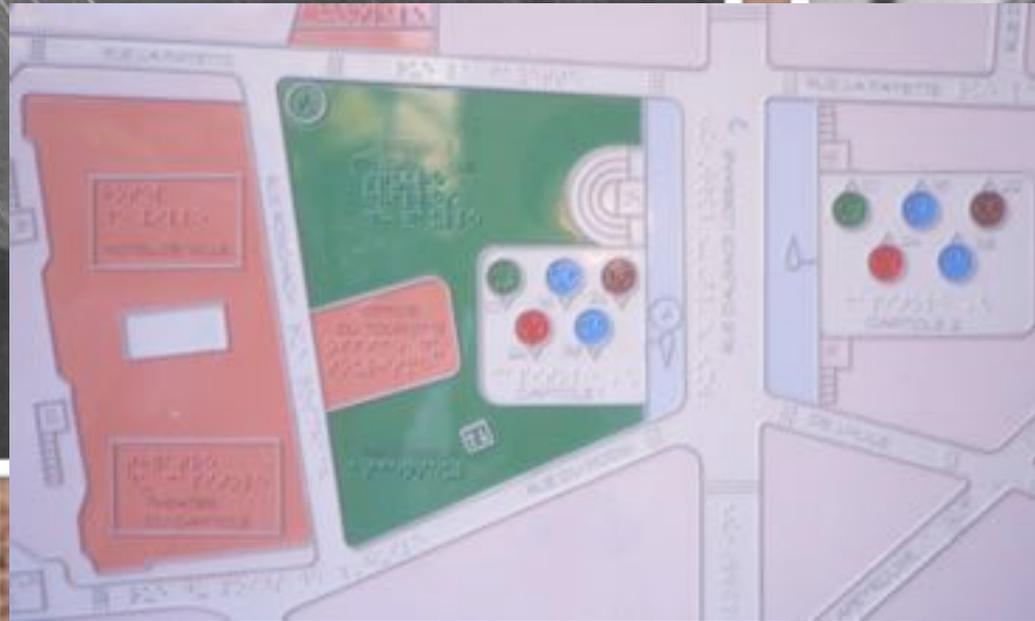
sombore

Paradoxe d'Olbers

cycle de la matière
expansion de l'Univers
architecture de l'Univers
matière noire
exoplanètes
énergie sombre
etc ...

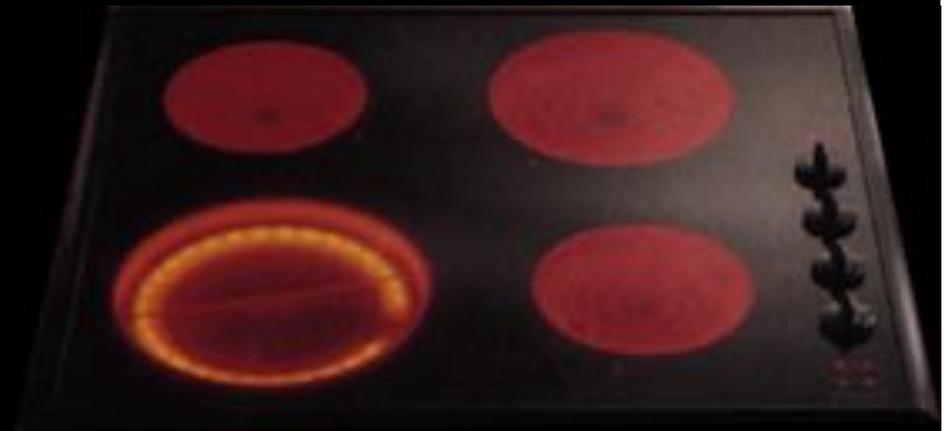
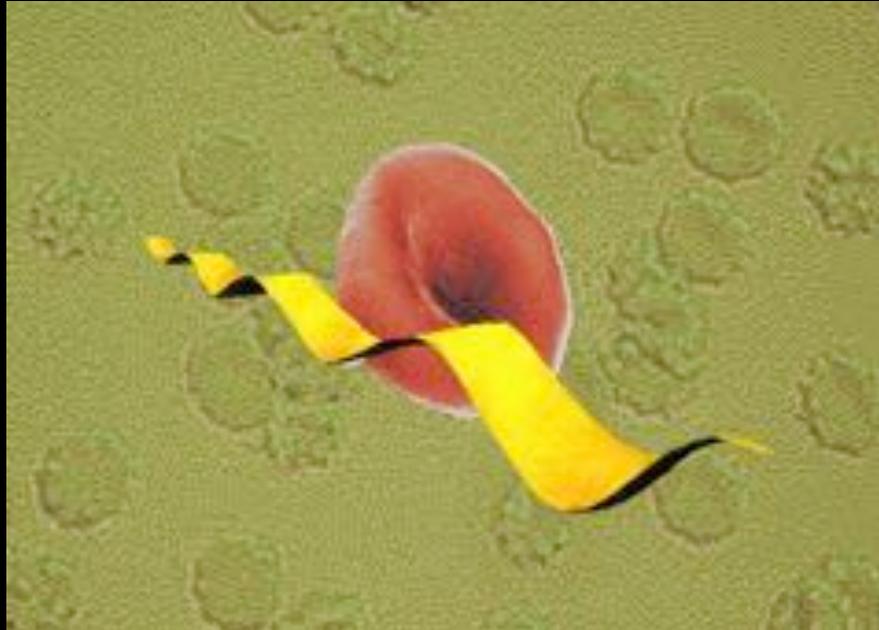


Comment imaginer l'invisible ?



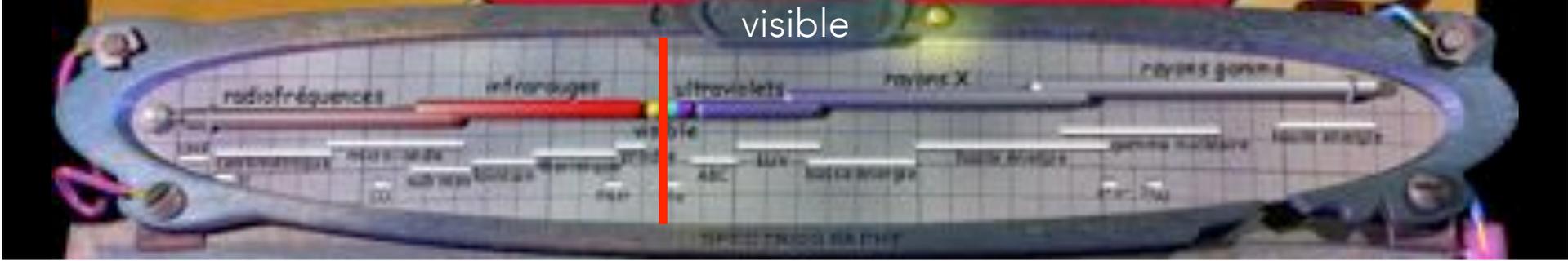
Plan de la station "Capitole" du Métro Toulousain

tableau de bord : infrarouge





visible





IR

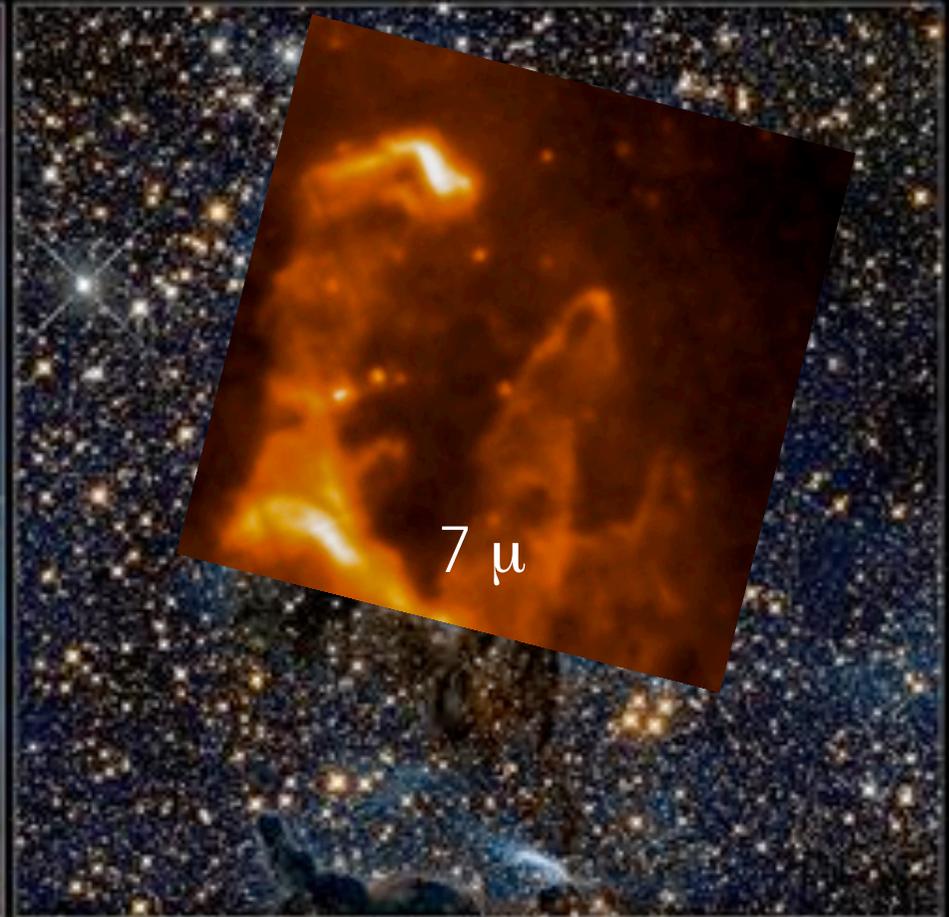


M16 – les "pilliers de la création"



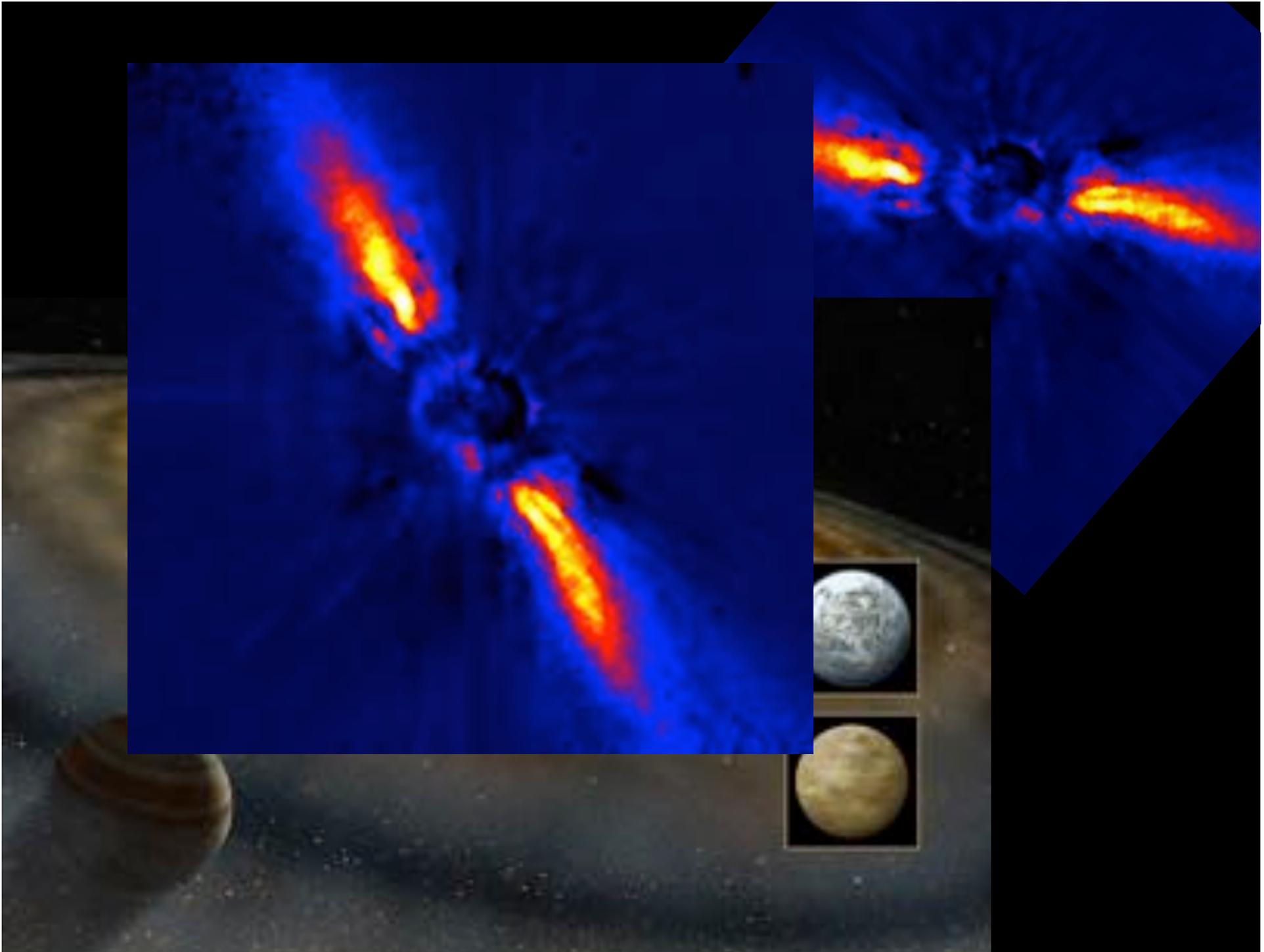
Visible - WFC3 - 2015

visible



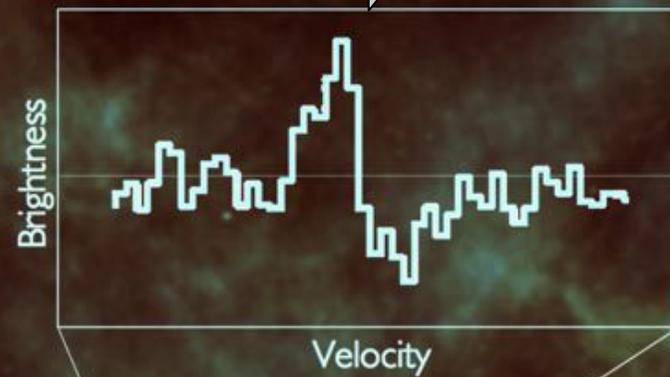
Infrared - WFC3 - 2015

infrarouge 1.1 et 1.6 μ



HERSCHEL
lancé en 2009

eau !



le noyau pre-stellaire "Lynds 1544"

Comment comprendre un ciel "oval" ?



les instruments de l'astronomie multi-longueur d'onde

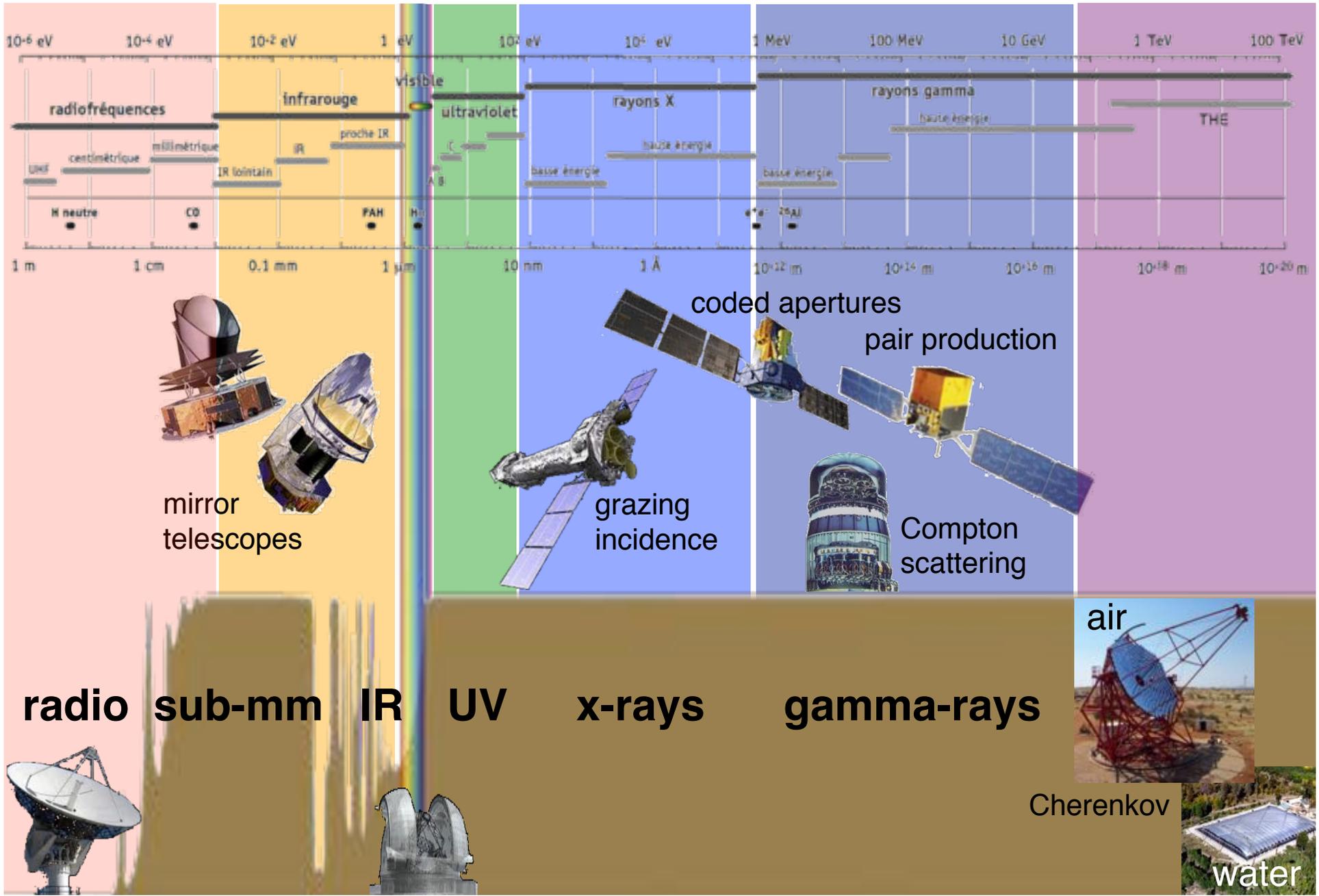
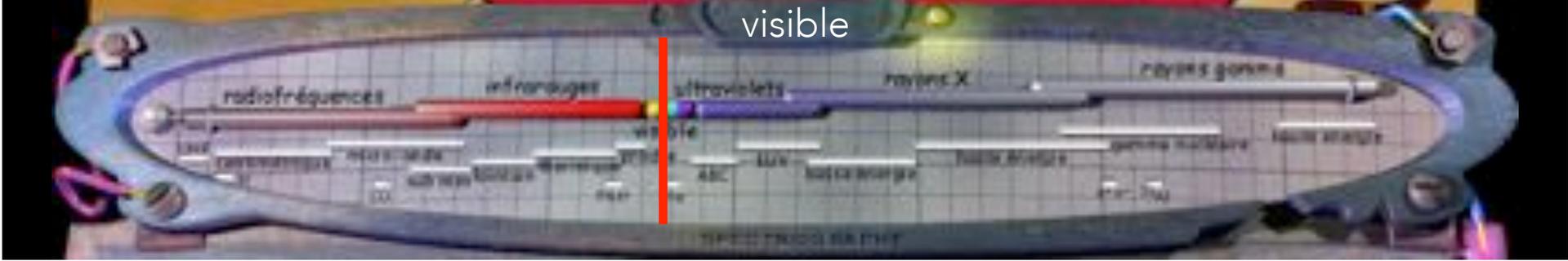


tableau de bord : micro-ondes



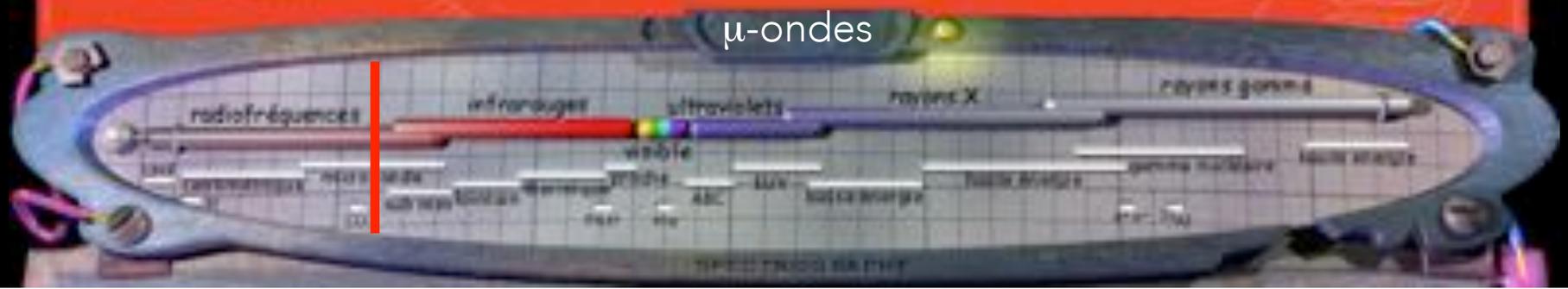


visible



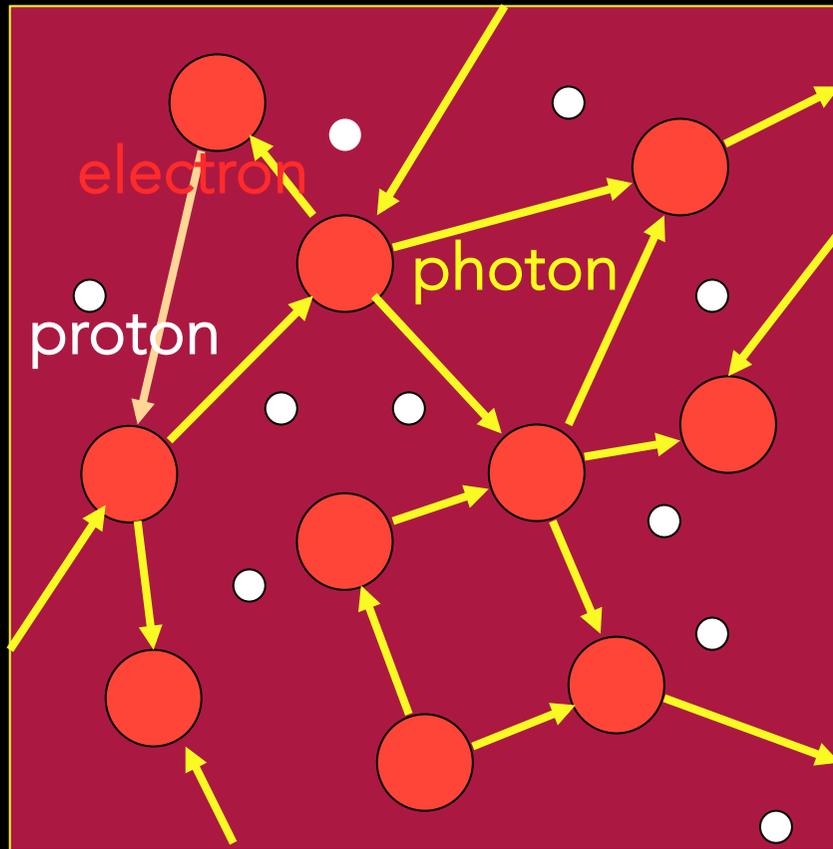


μ -ondes



380'000 ans après le Big Bang lumière et matière "divorcent"

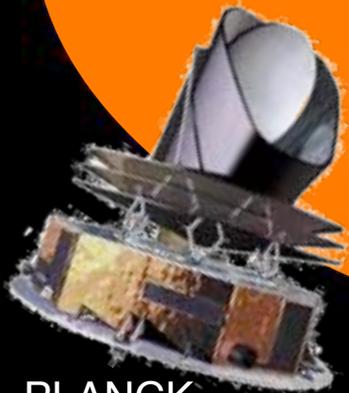
avant



l'Univers devient transparent – les atomes (H) naissent



fond cosmologique
age précis de l'Univers
geometrie de l'Univers

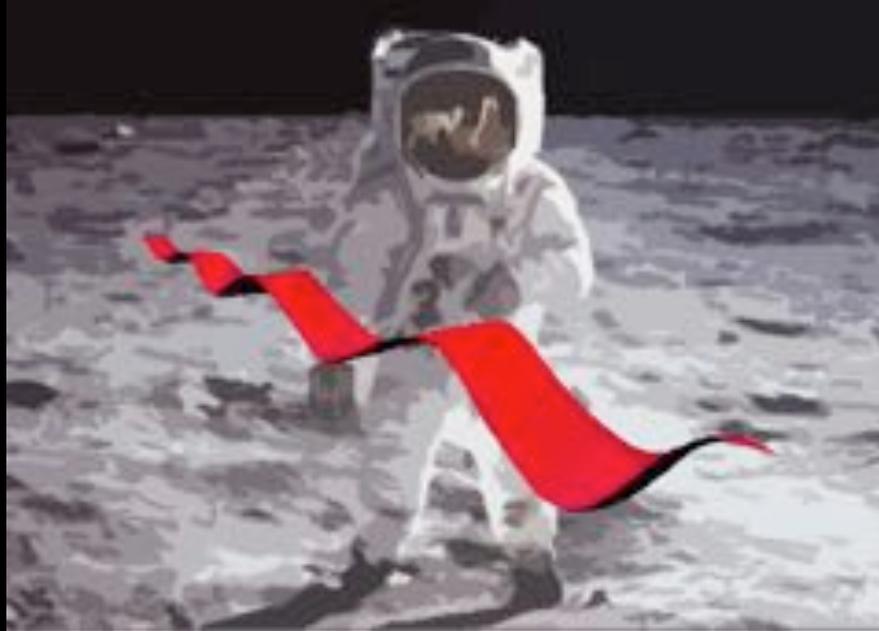


PLANCK

=> L'Univers a 13.8 Milliards d'années

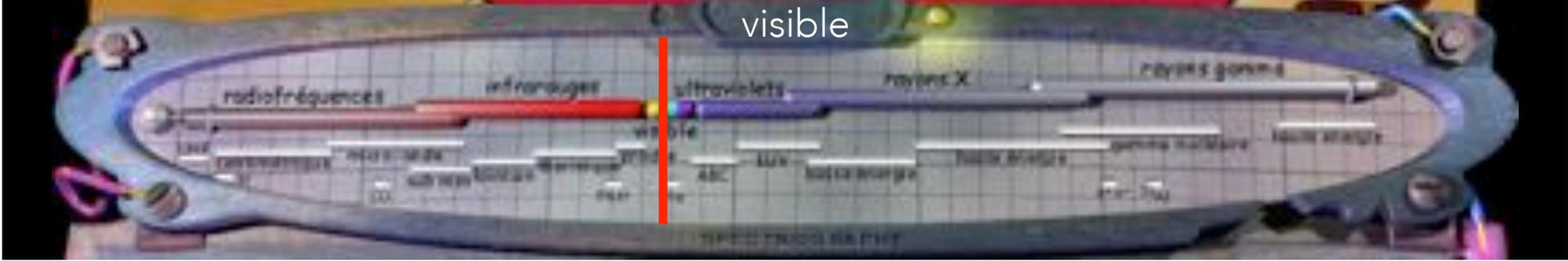


tableau de bord : ondes radio





visible





radio



le reste de
supernova de la Vela

digitized sky survey

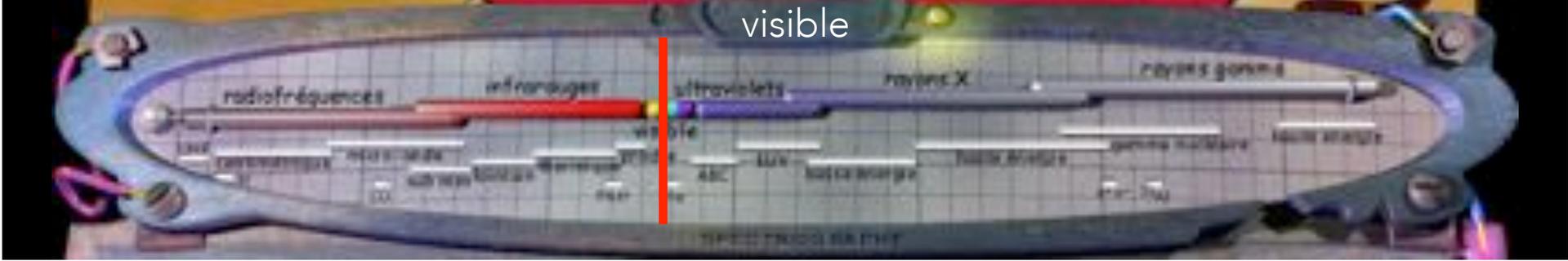


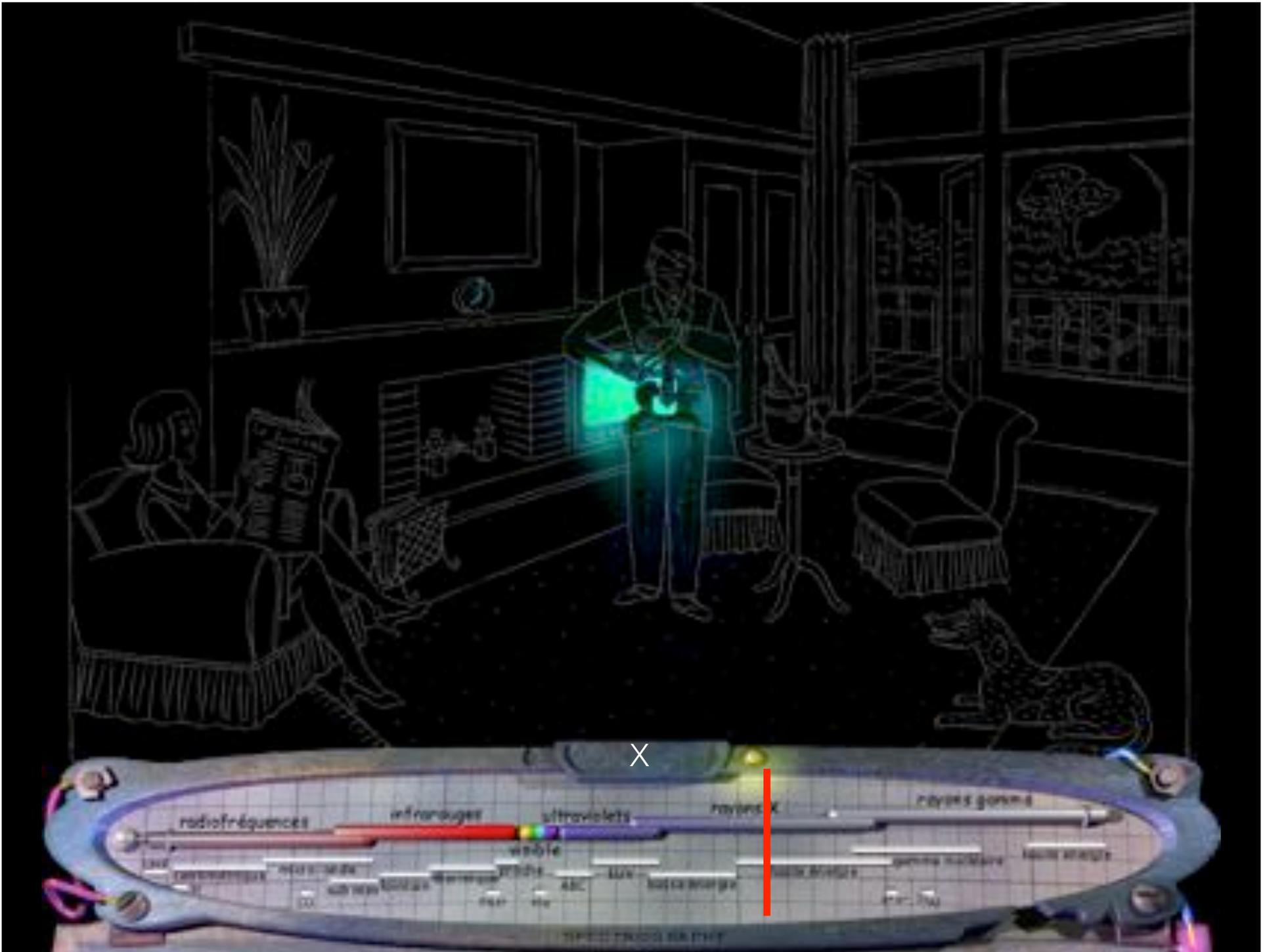
tableau de bord : rayons X





visible





le reste de
supernova de la Vela

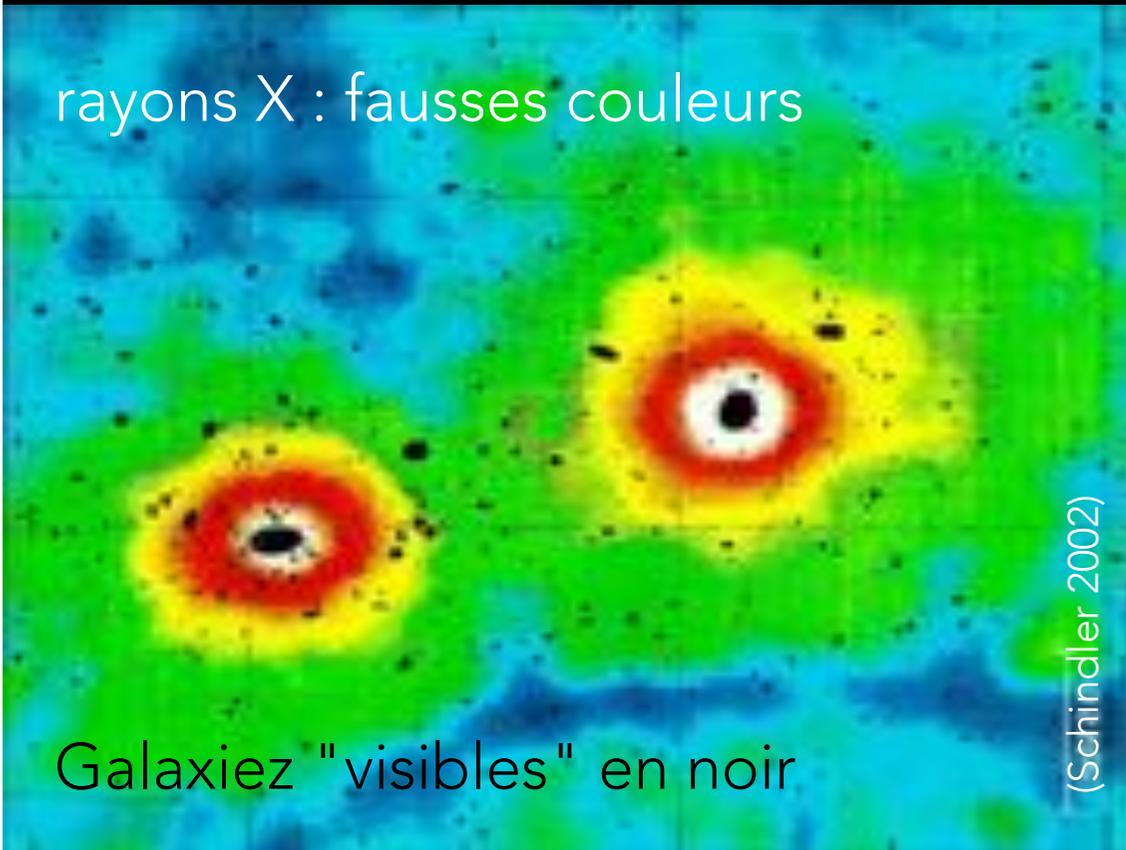
digitized sky survey



Abell 3528
"merger"
de deux amas de galaxies



rayons X : fausses couleurs



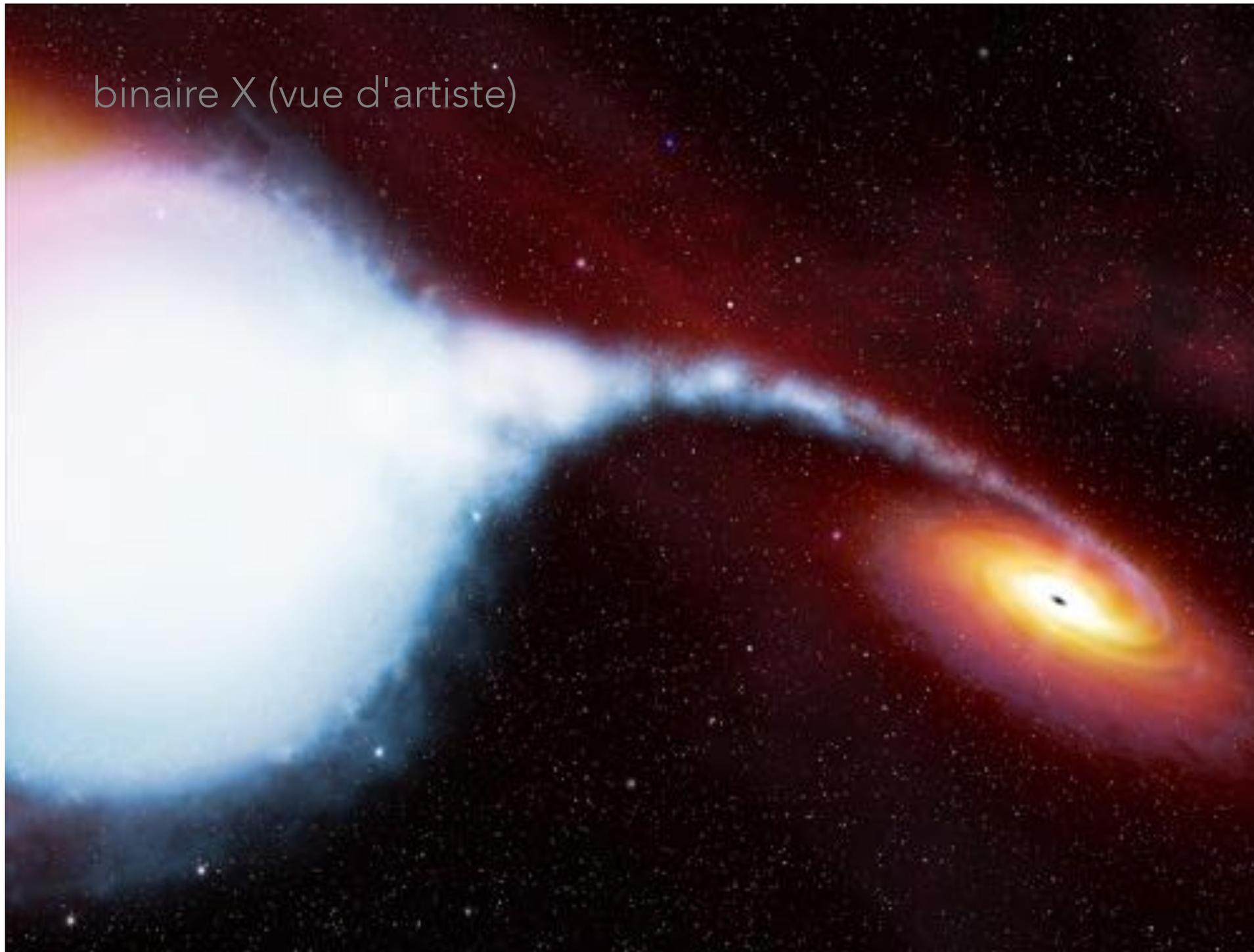
Galaxiez "visibles" en noir

(Schindler 2002)

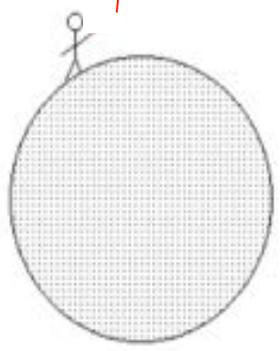
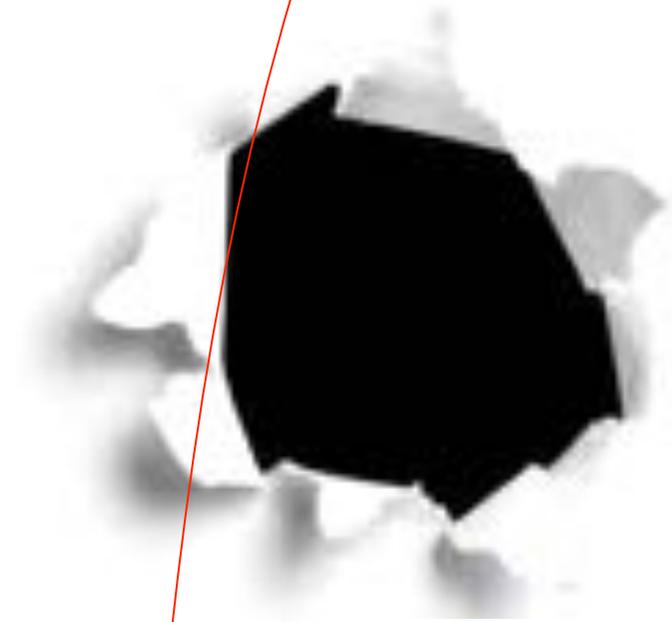


ROSAT

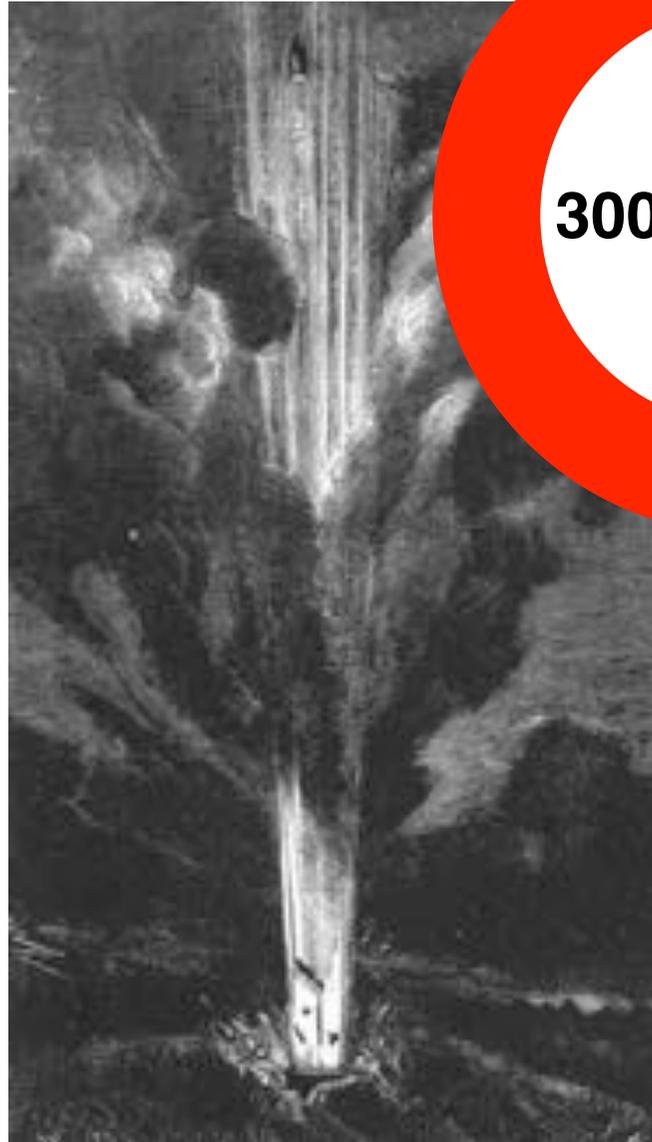
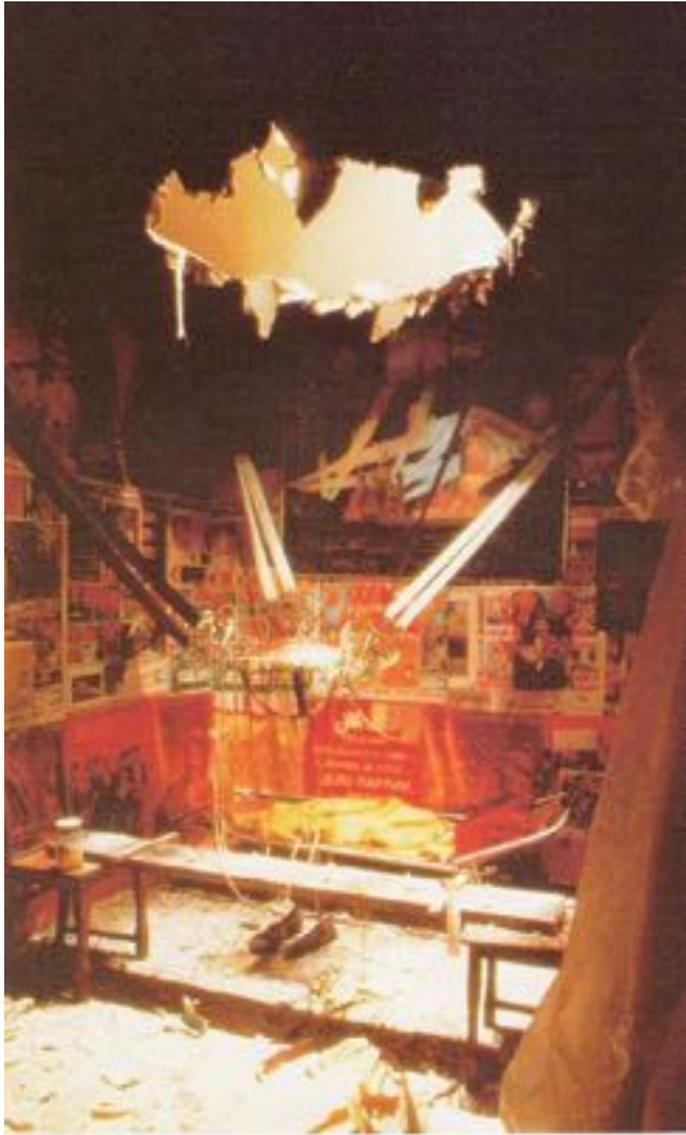
binaire X (vue d'artiste)



Qu'est ce qu'un trou noir ?



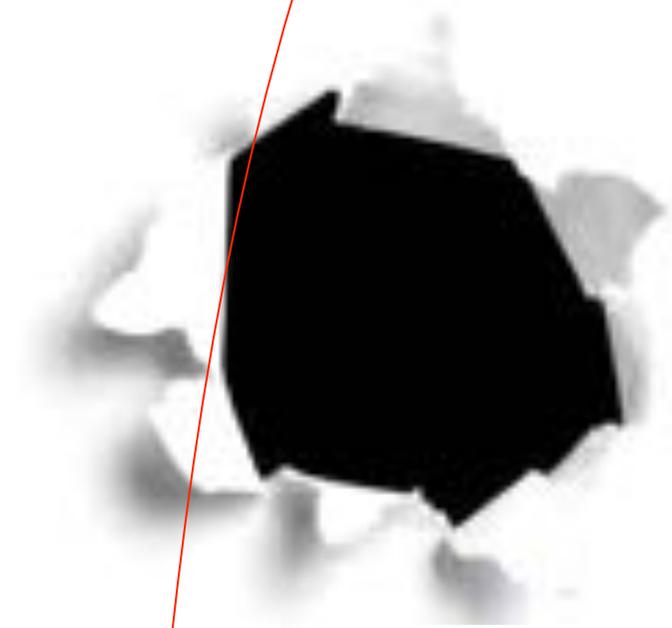
vitesse de liberation



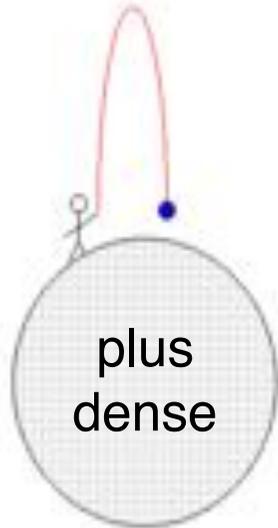
300'000 km/s

pour se liberer de la Terre : 11 km/sec

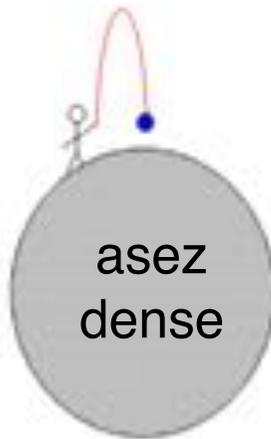
Qu'est ce qu'un trou noir ?



peu
dense



plus
dense



assez
dense



très
dense



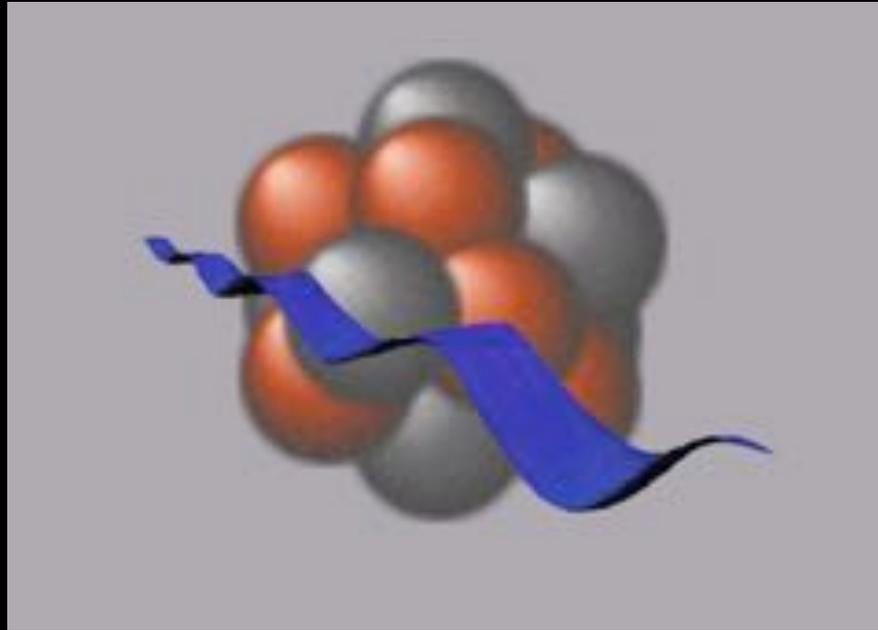
très
très
dense

Swift J1644+57 : une nouvelle source de haute énergie dans le "dragon"



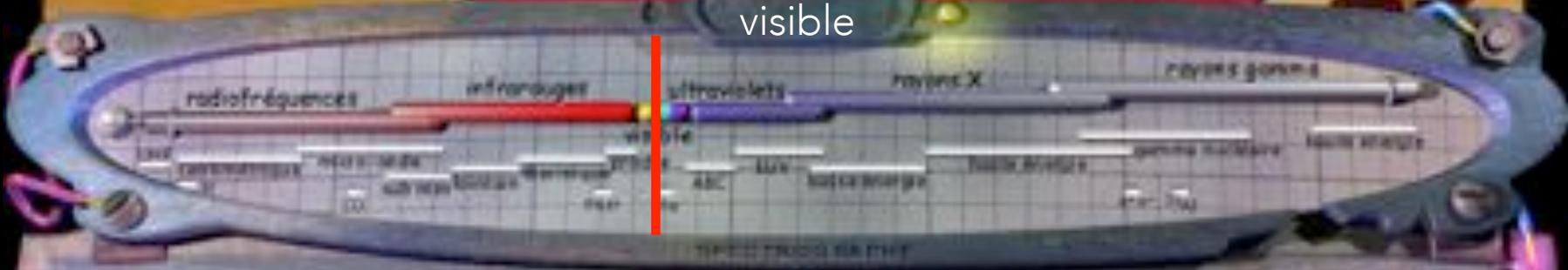
animation

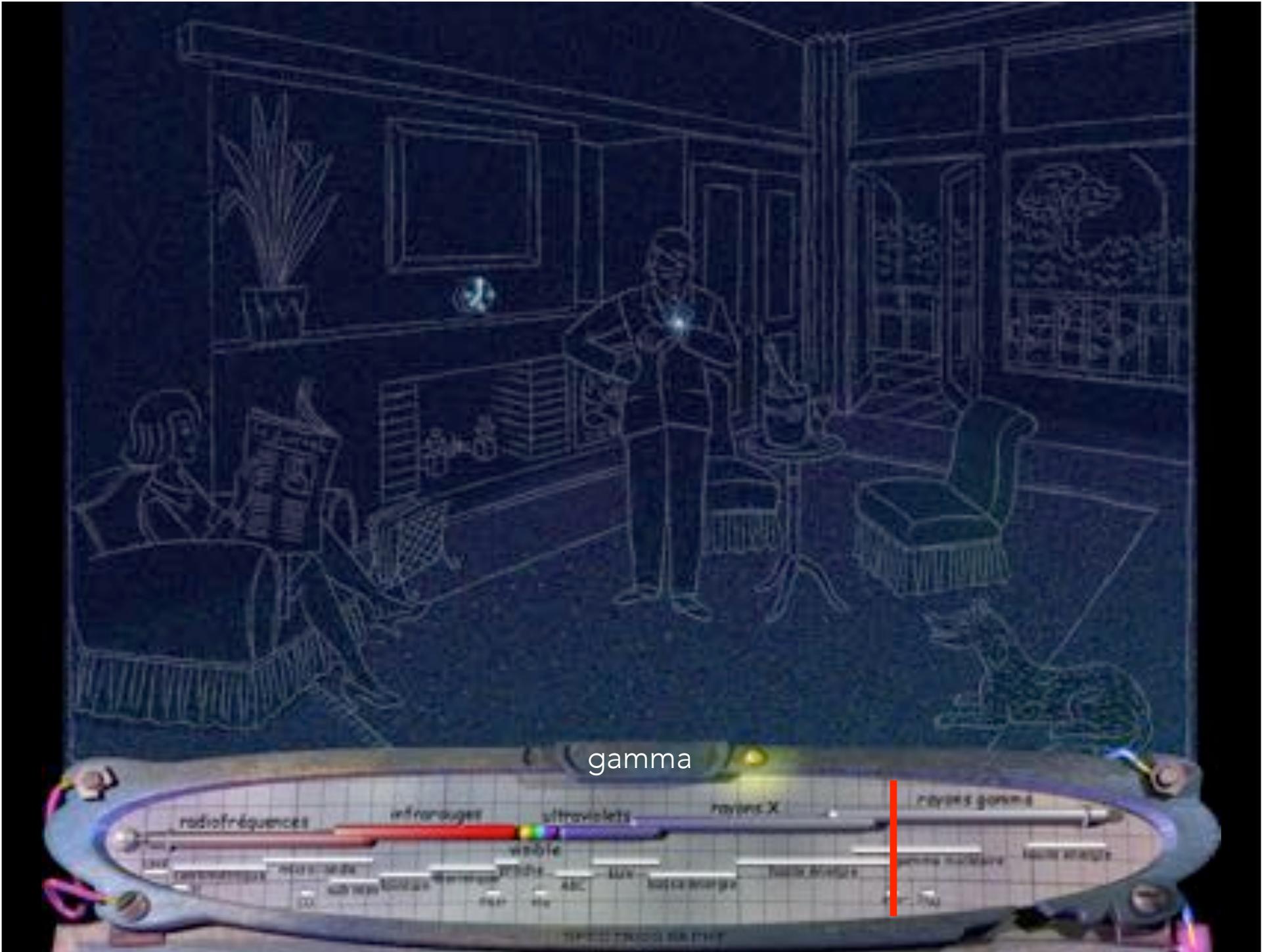
tableau de bord : rayons gamma





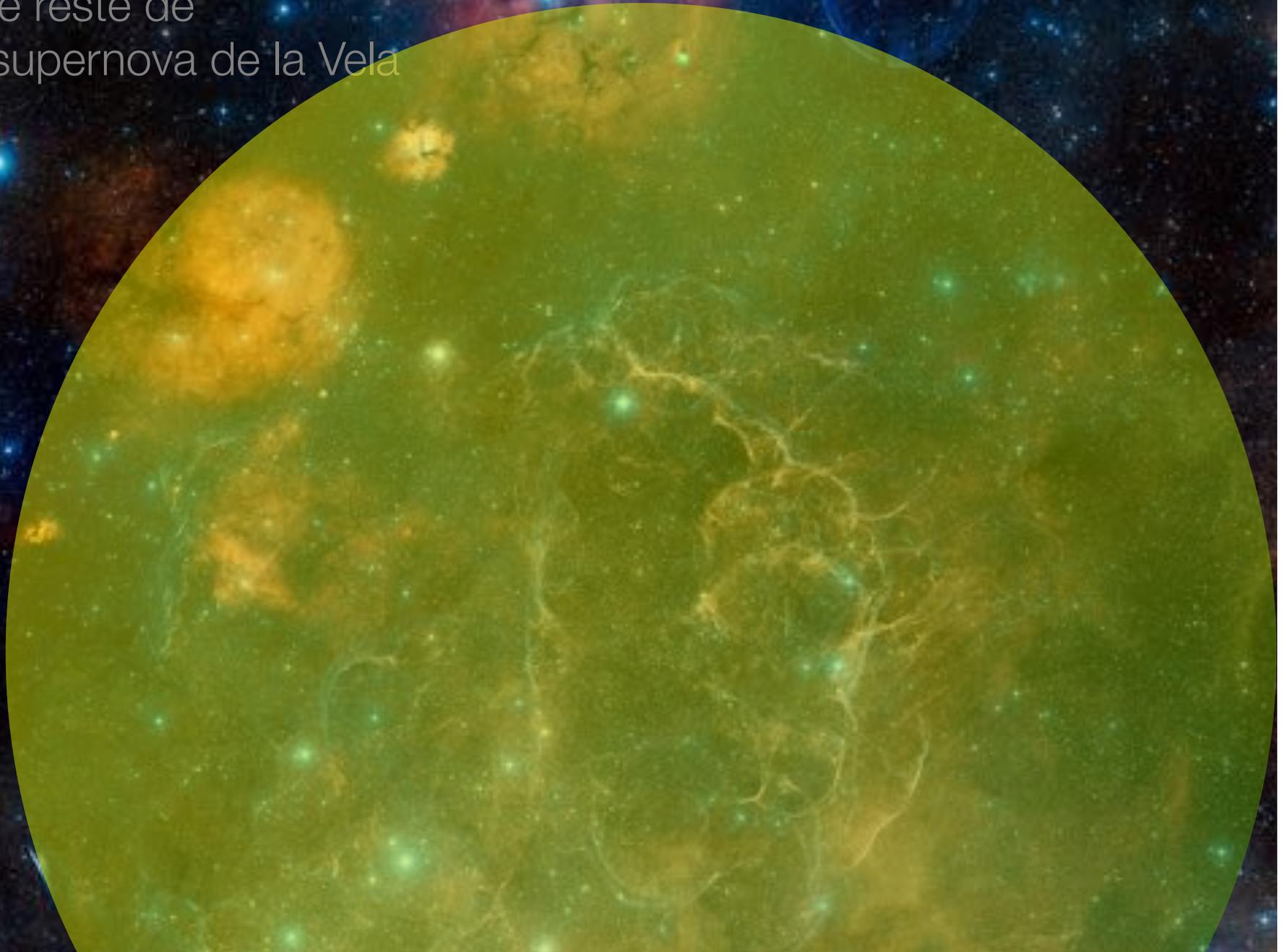
visible





le reste de
supernova de la Vela

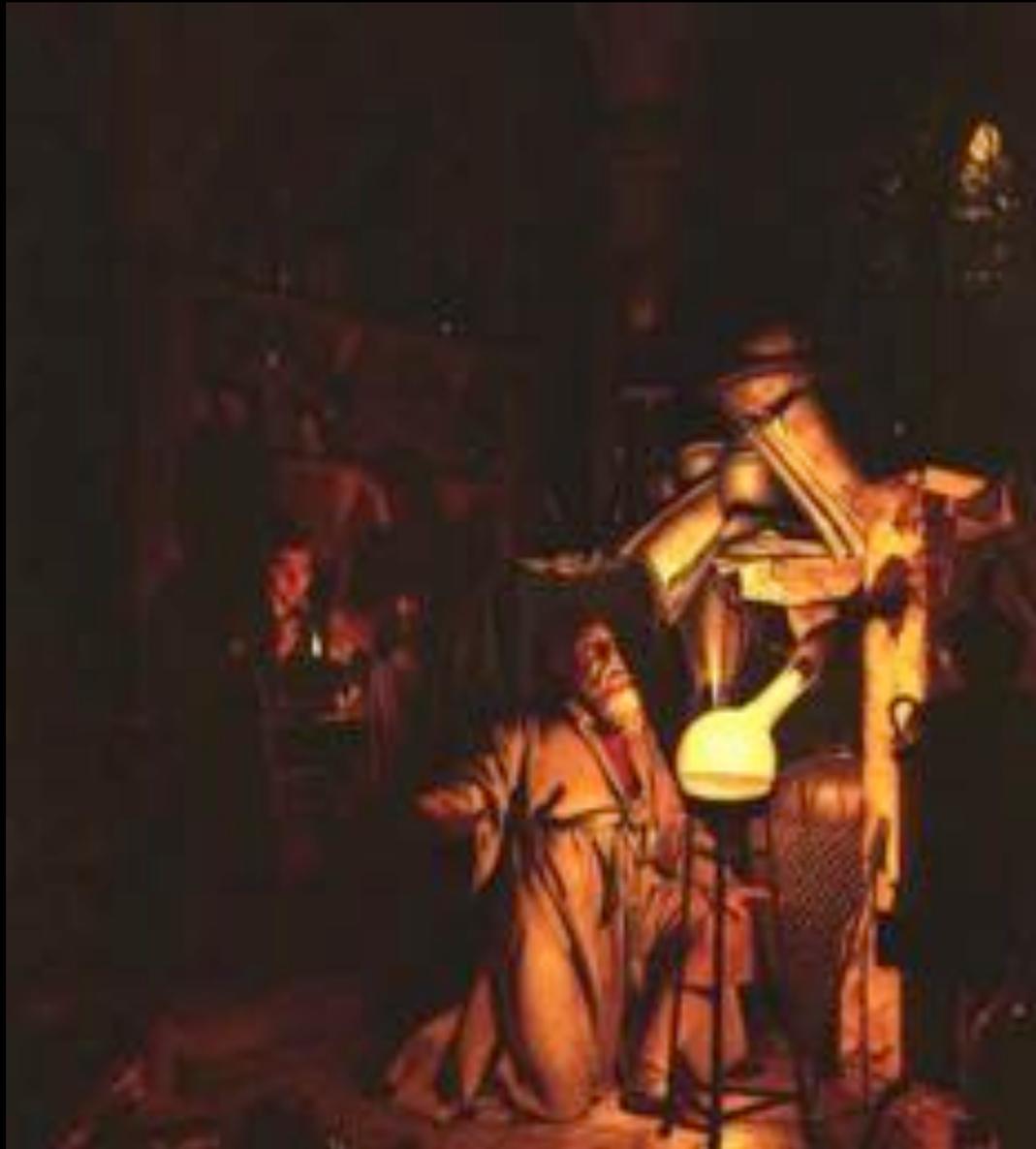
digitized sky survey



les éléments chimiques



évolution stellaire - le cycle de la matière

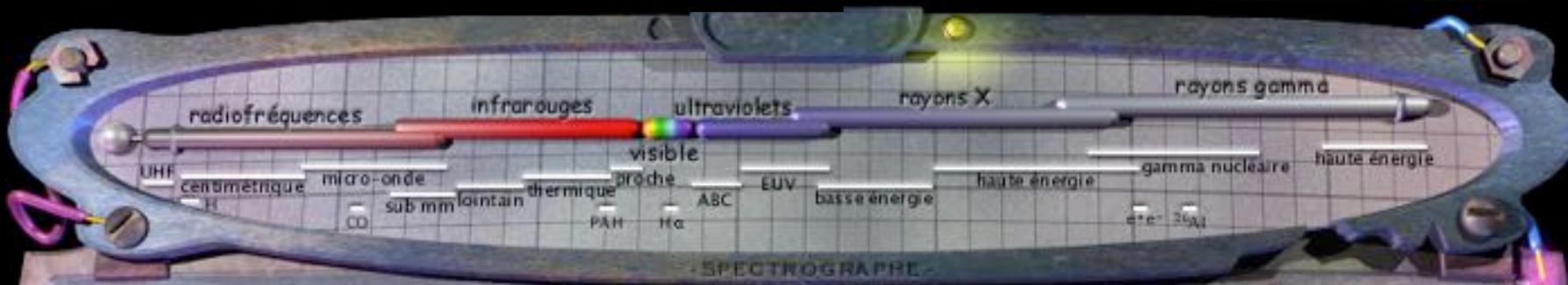
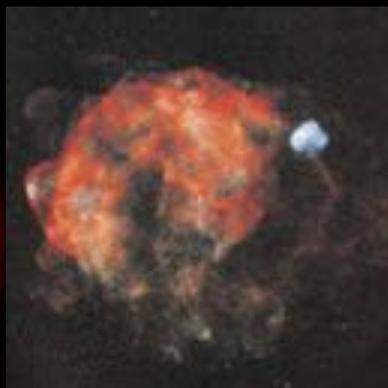
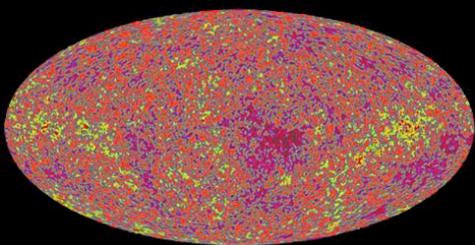
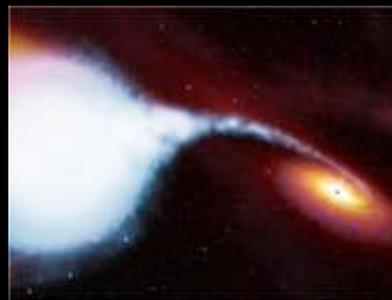
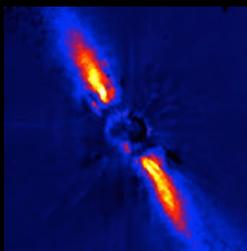
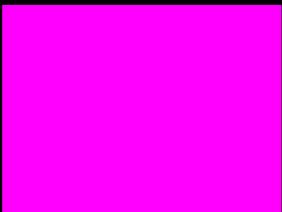


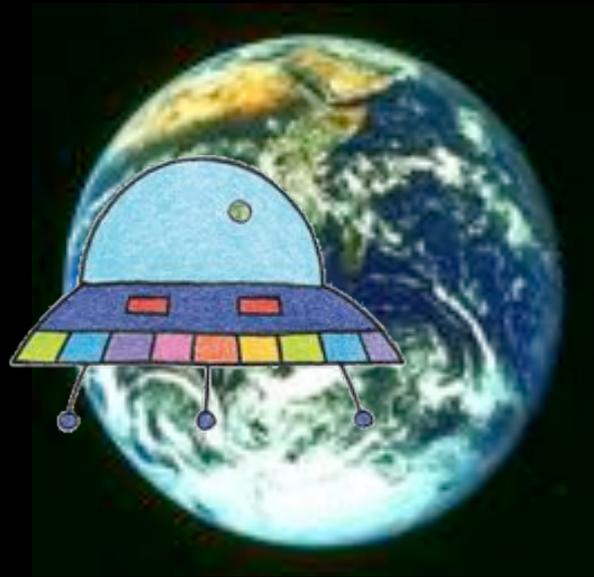
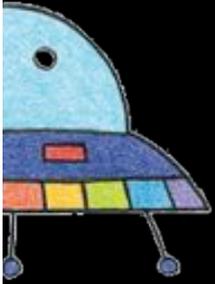
radioactivité
sursauts gamma
annihilation e^-e^+
magnetars ...



INTEGRAL

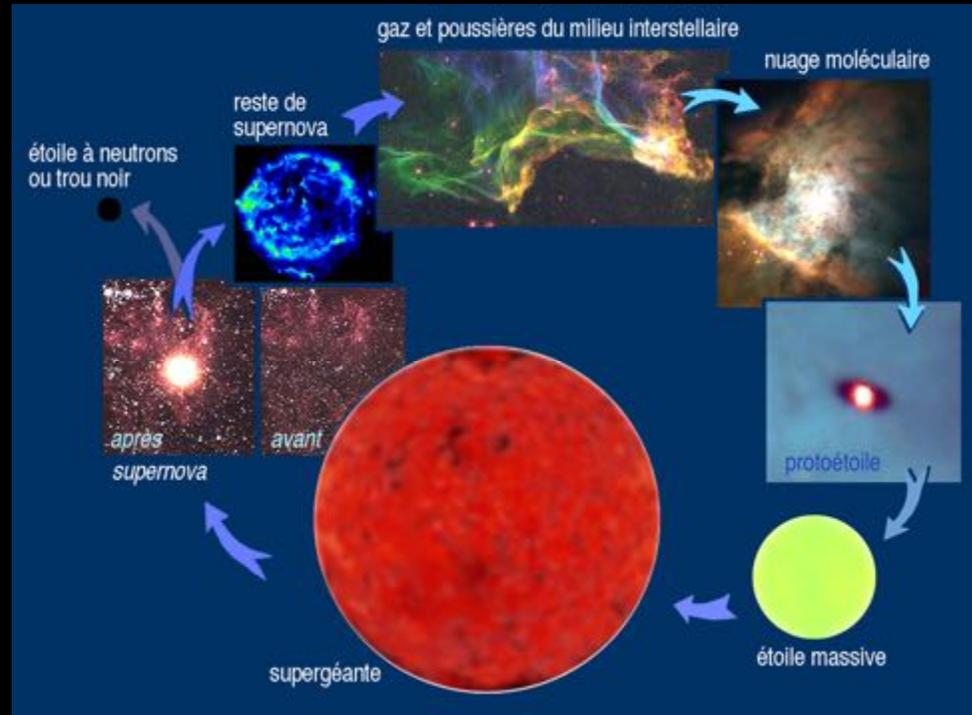






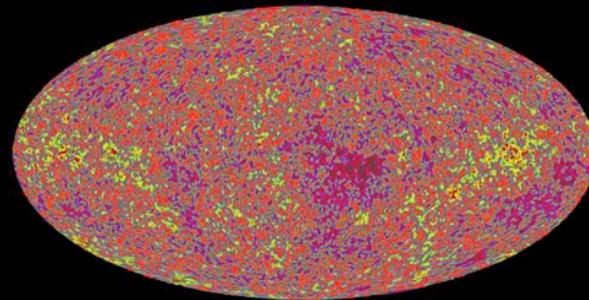


- les éléments qui nous constituent sont produites dans les étoiles



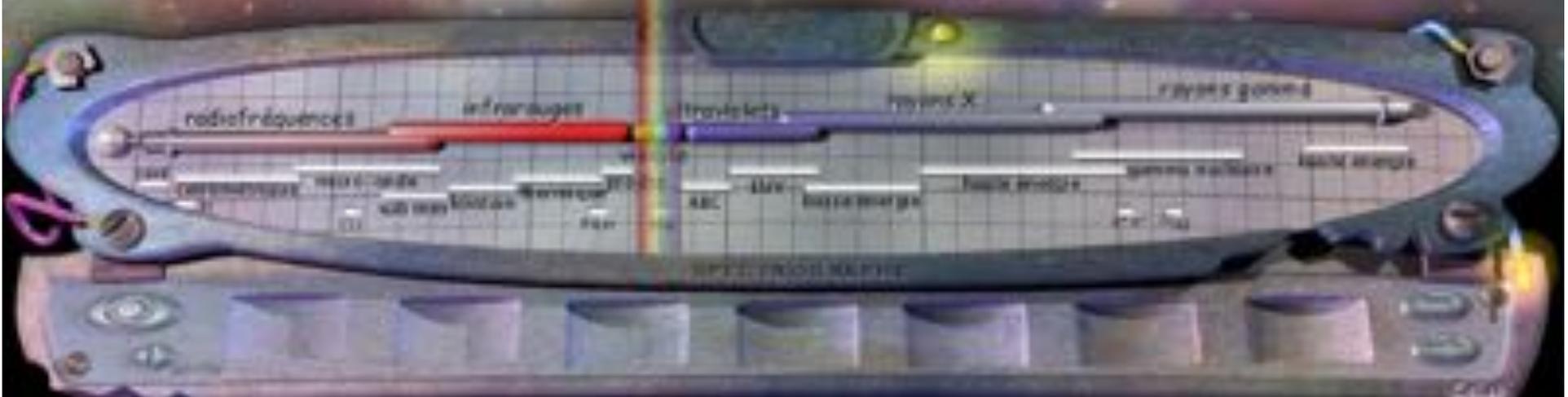
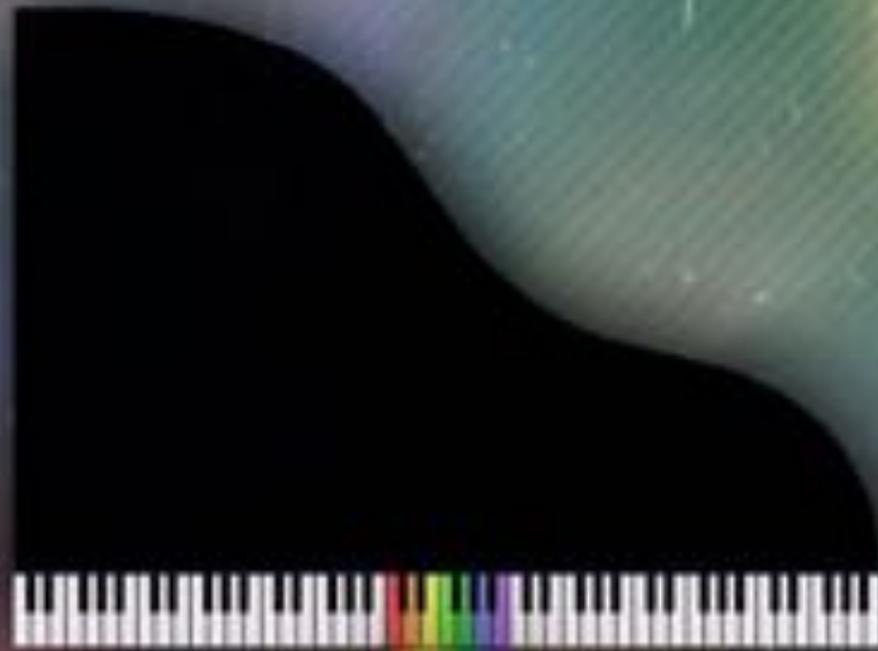
le cycle de la matière

- l'Univers a un début



comment imaginer l'invisible ?

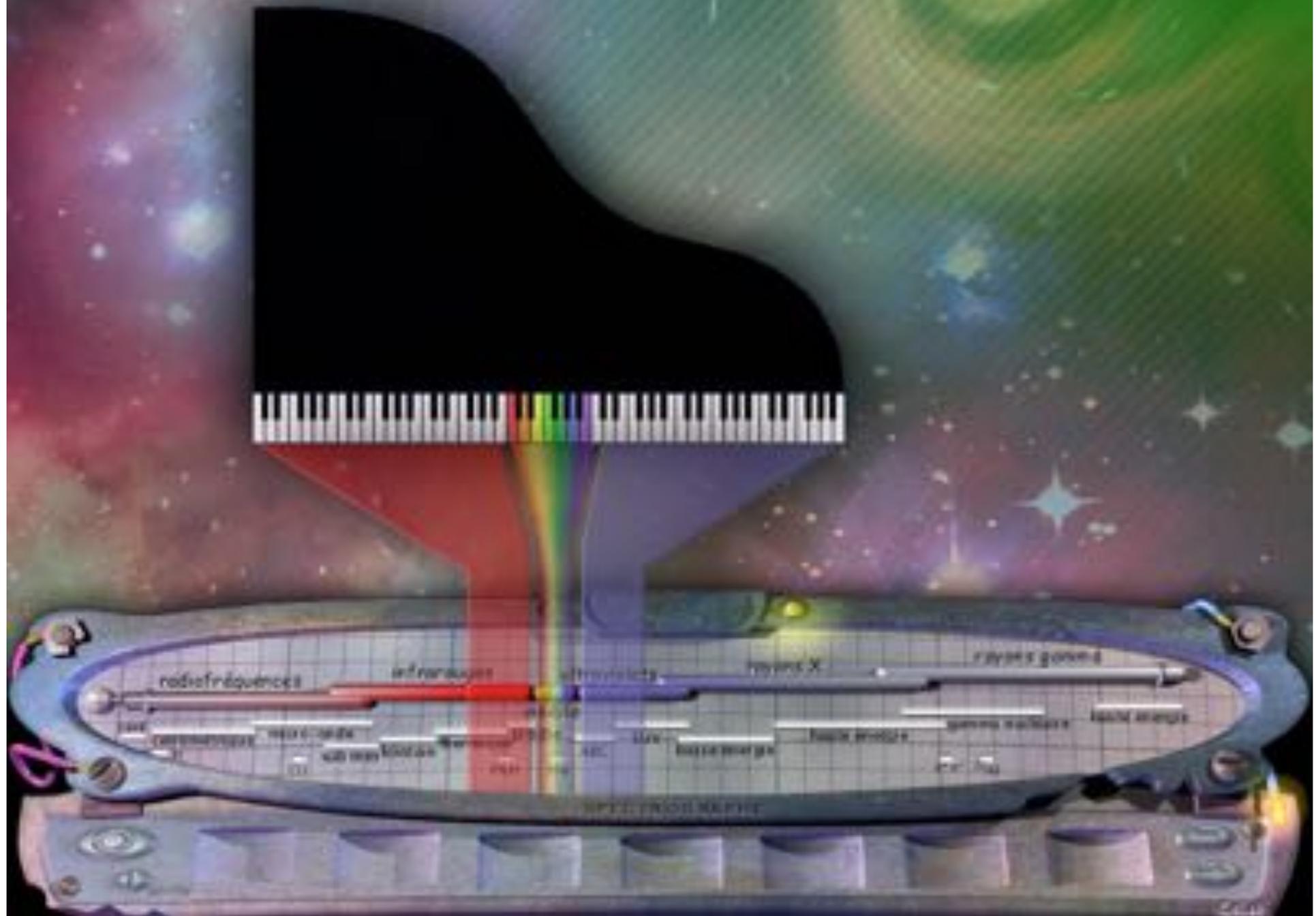
Créa



Créa

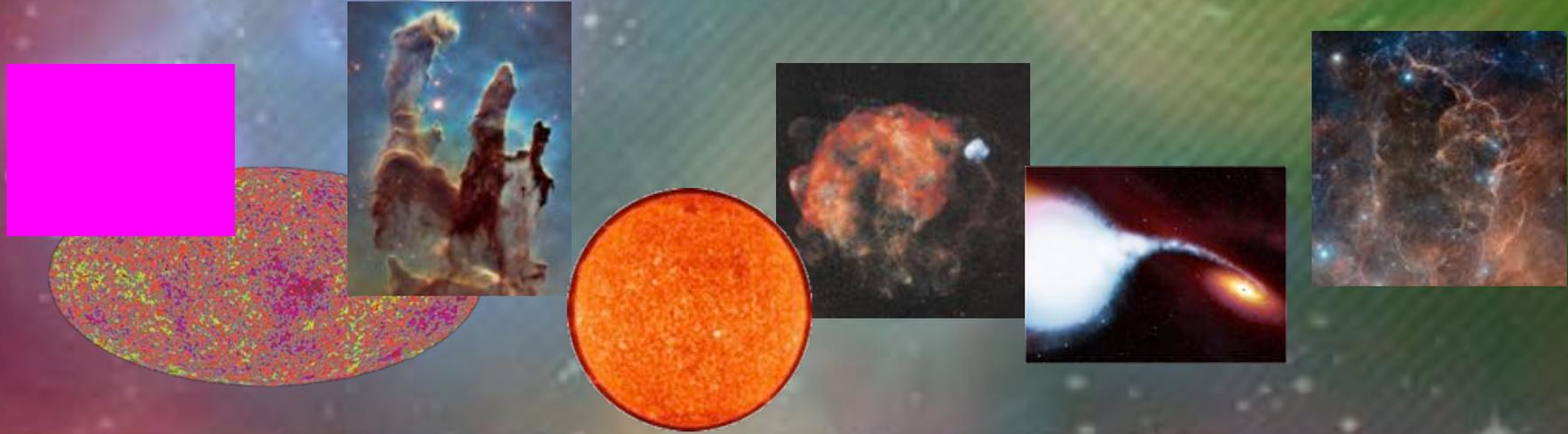
comment imaginer l'invisible ?

Créa



Créa

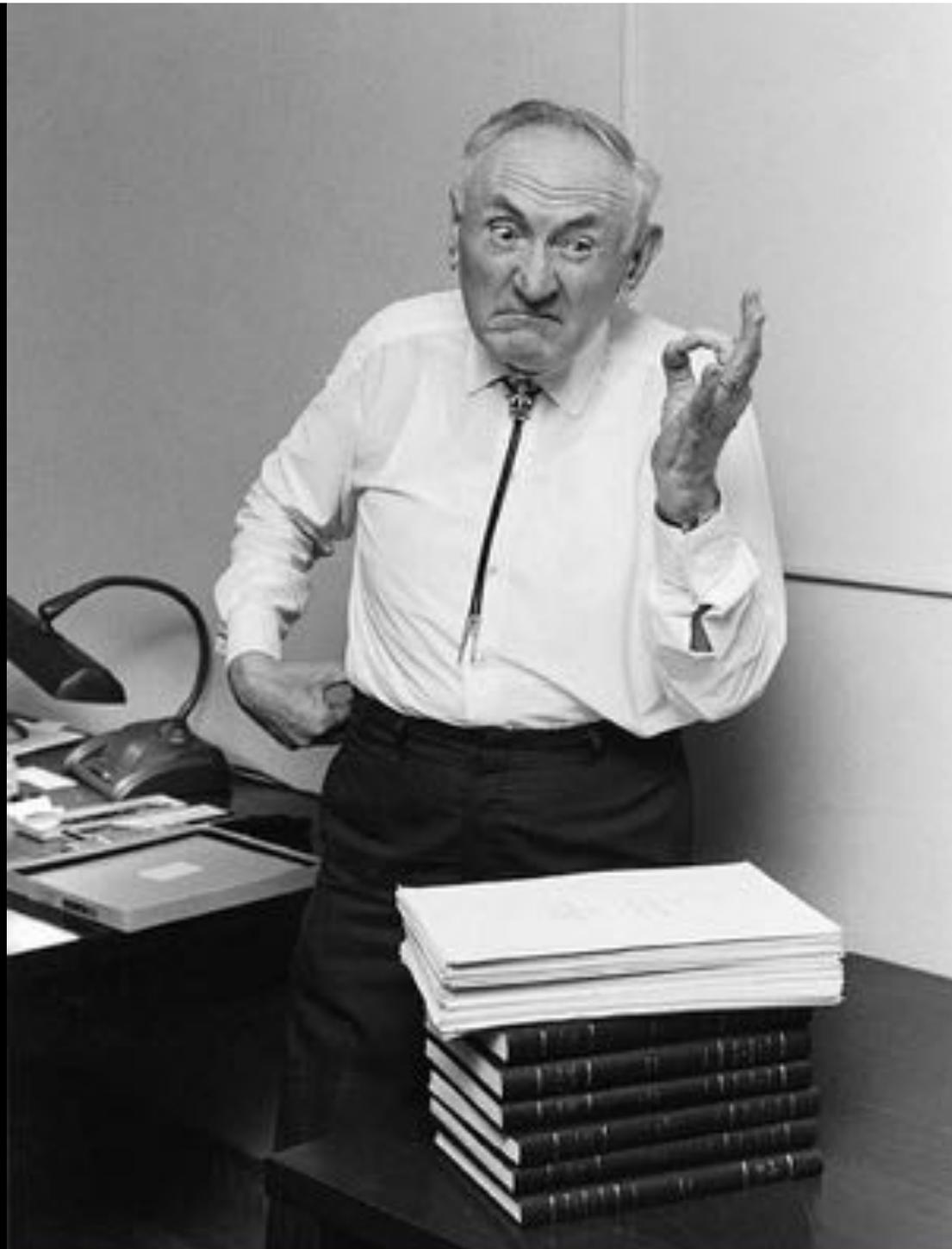
comment imaginer l'invisible ?



4 % de ce qu'il y a dans l'Univers

Qu'est ce que la matière noire ?



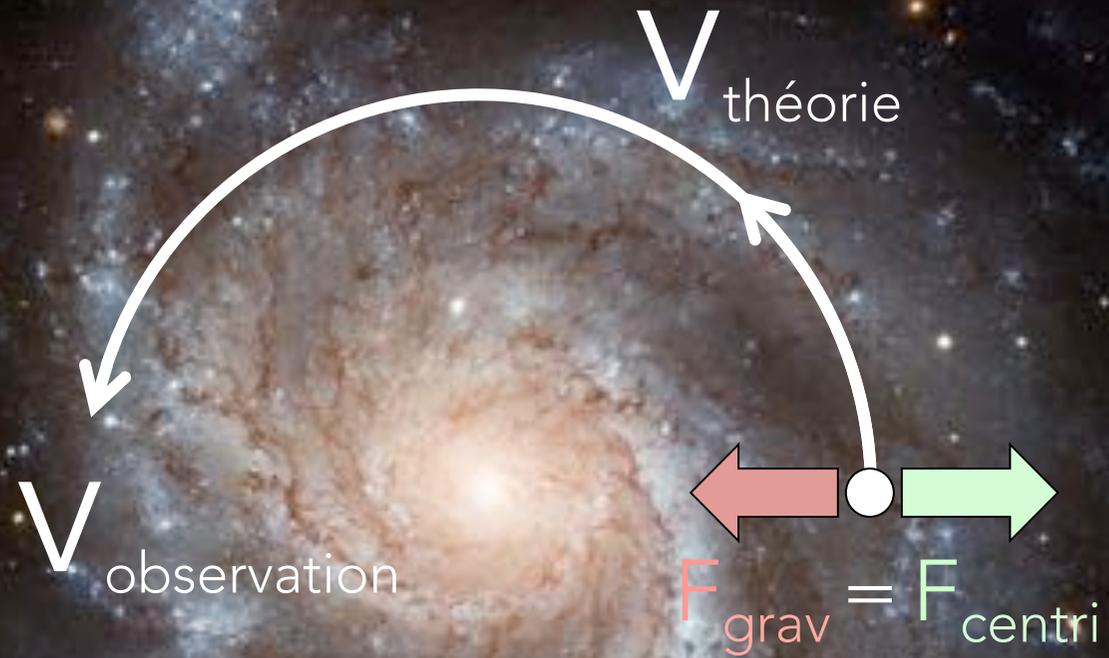


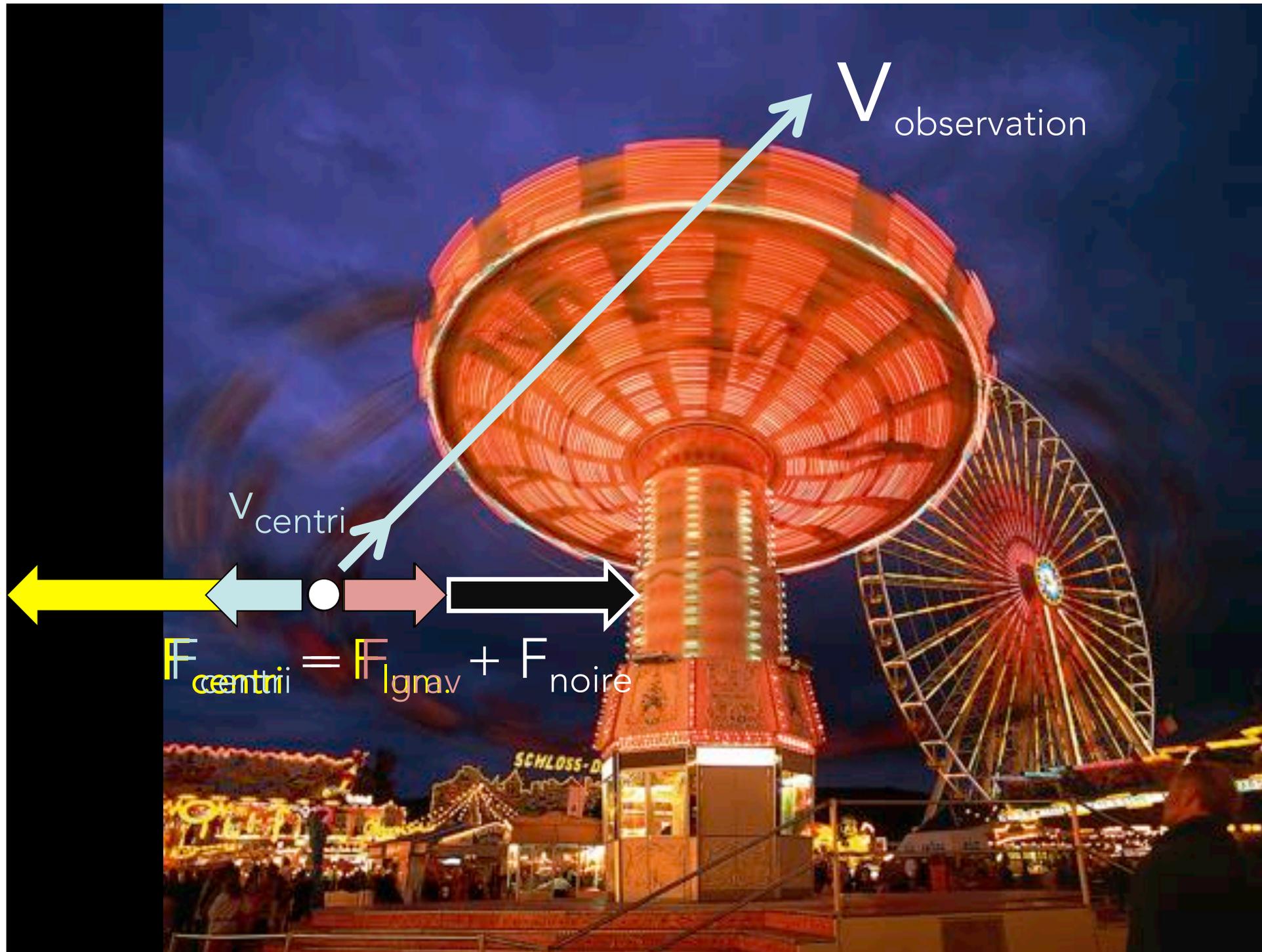
Fritz Zwicky
(1898-1974)

l'amas de Galaxies Coma

Zwicky étudie la distribution des vitesses des Galaxies
sa conclusion : il faut $M_{\text{gravitation}} \gg M_{\text{visible}}$
sinon les galaxies seraient éparpillées par leur vitesse.

ça tourne ... trop vite





lentilles gravitationnelles

gravitational lensing par des amas de galaxies => $\Omega_{\text{matière}} \approx 0.3$



Galaxy Cluster Abell 2218

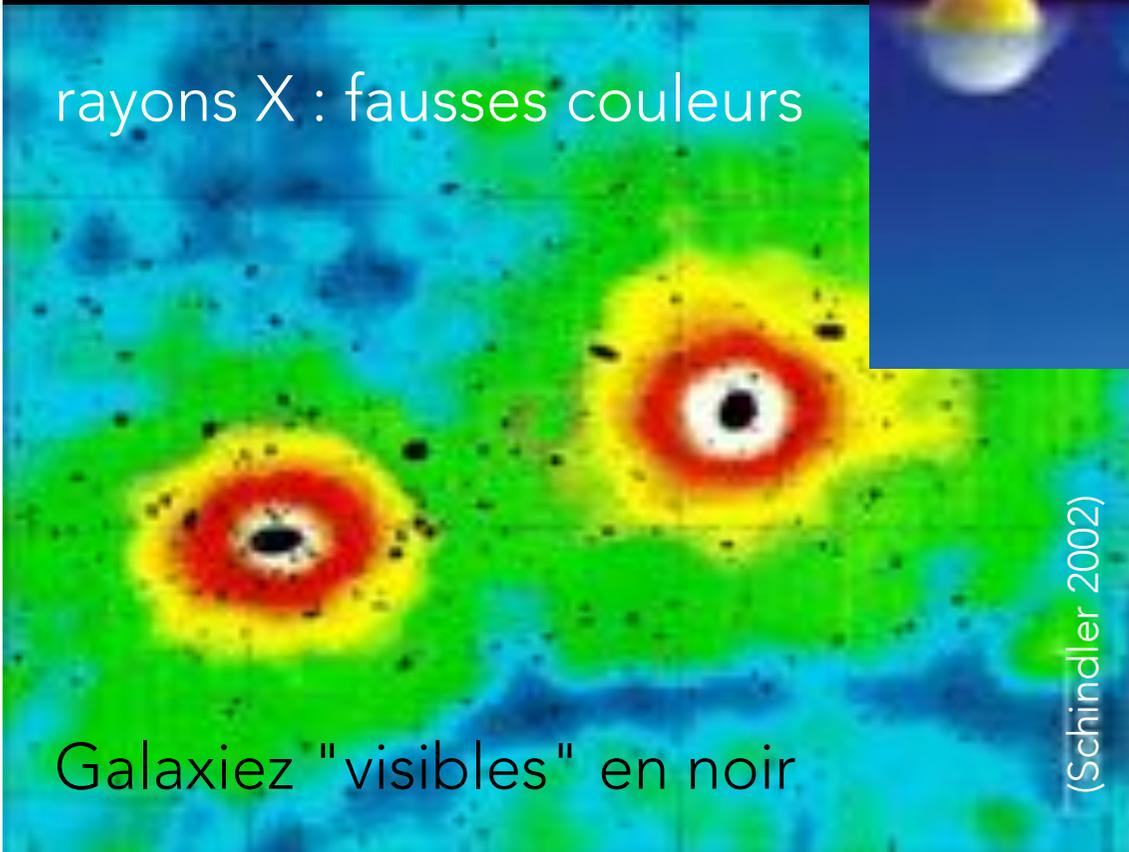
HST • WFPC2

NASA, A. Fruchter and the ERO Team (STScI) • STScI-PRC00-08

Abell 3528



rayons X : fausses couleurs



Galaxies "visibles" en noir

(Schindler 2002)



les "puits de potentiel"
de la matière noire
sont remplis de gaz chaud

Quelle est la nature de la matière noire?

De la matière « normale » mais obscure?

- ~~poussières?~~
- ~~planètes, astéroïdes?~~
- étoiles noires, brunes, blanches?
- ~~trous noirs stellaires?~~
- ~~trous noirs géants?~~

Machos

~~Neutrinos?~~

Axions

Un nouveau type de particule élémentaire stable?

- très légère et très abondante?
- très lourde et très rare?

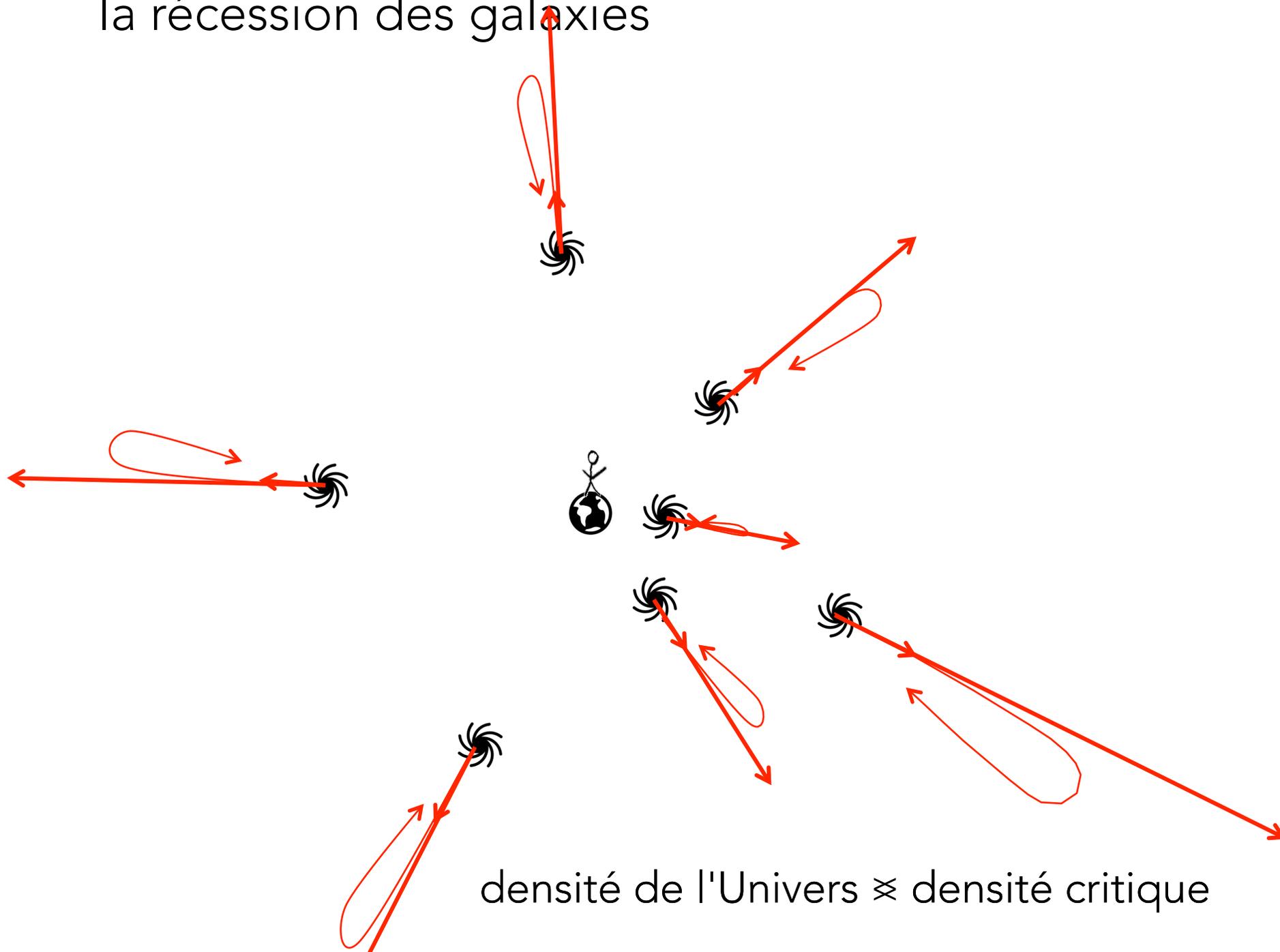
Wimps

~~Gravitation modifiée~~

Qu'est ce que l'énergie noire ?



la récession des galaxies

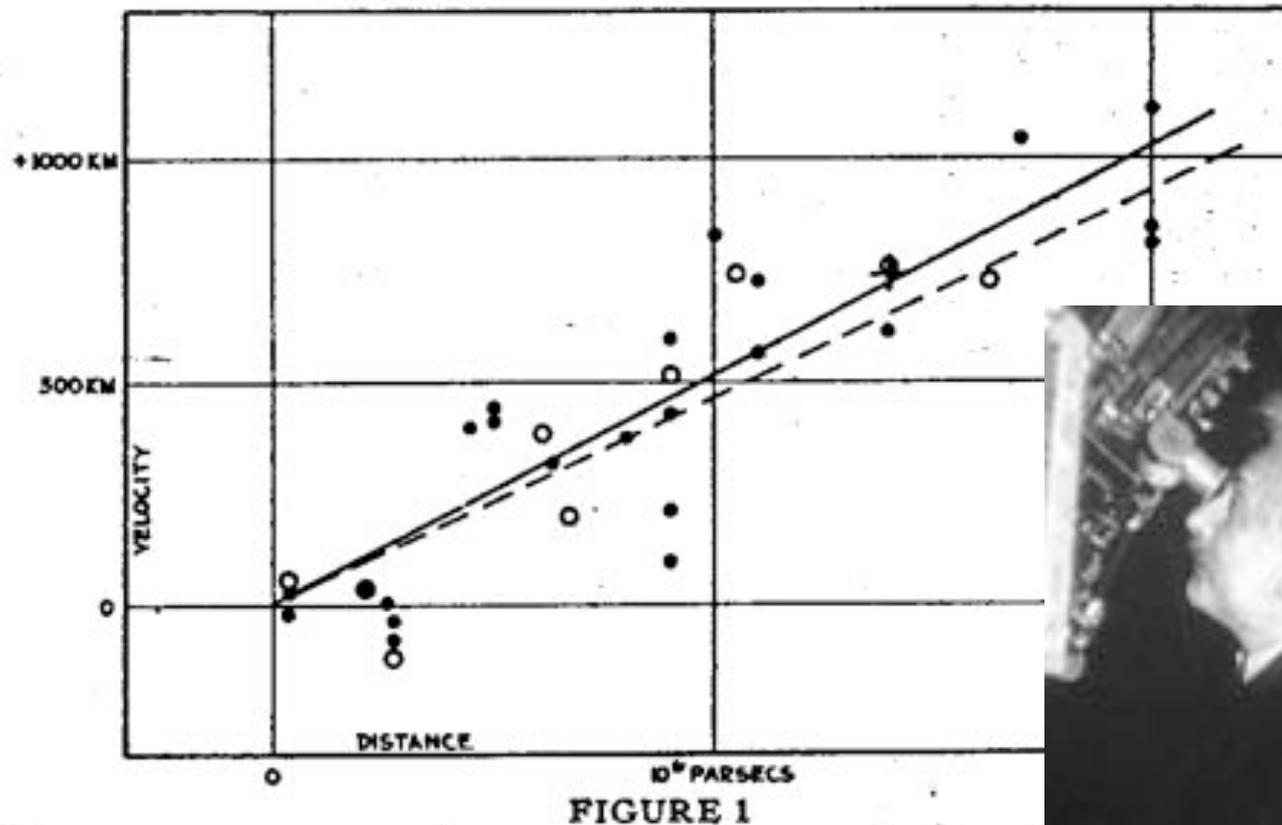


densité de l'Univers \approx densité critique

L'expansion de l'Univers : Le Diagramme de Hubble

A RELATION BETWEEN DISTANCE AND RADIAL VELOCITY
AMONG EXTRA-GALACTIC NEBULAE

By Edwin Hubble



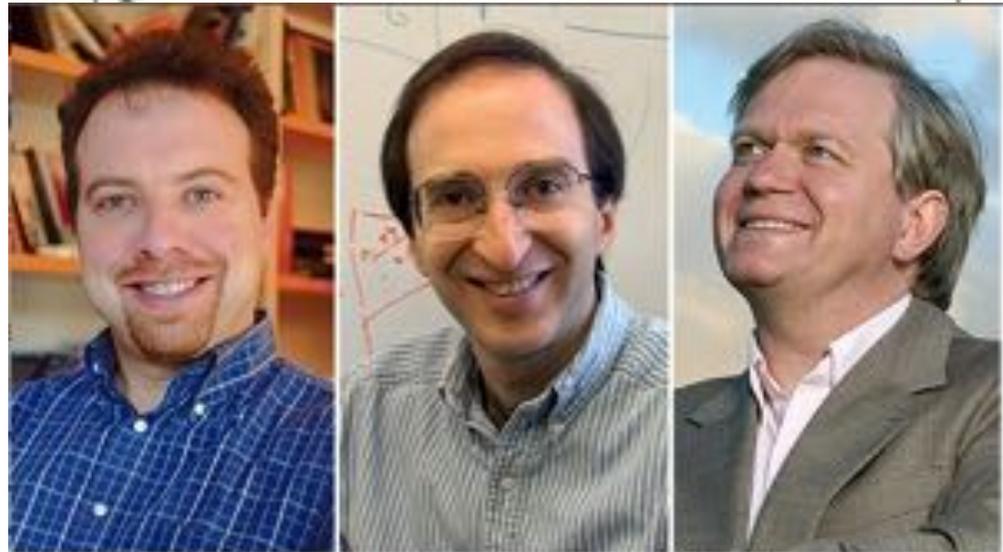
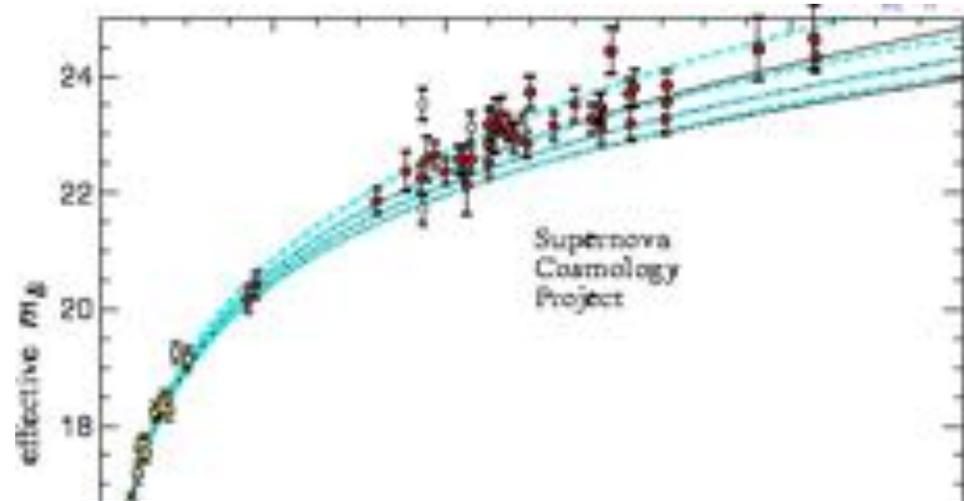
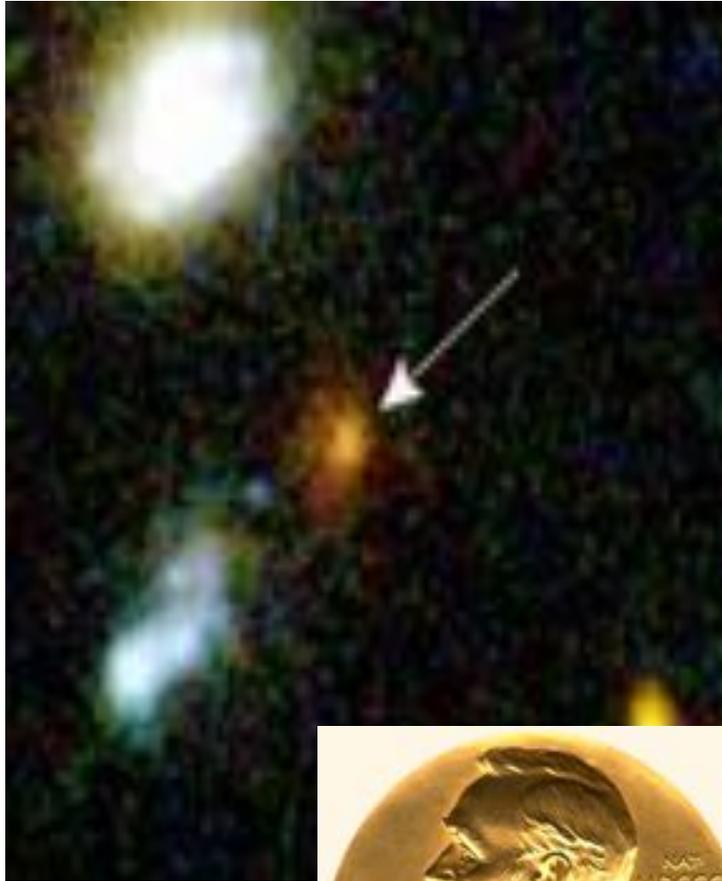
from the Proceedings of the National Academy of Sciences Volume 15 : March 15, 1929 : Number 3

Qu'est ce que l'énergie noire ?



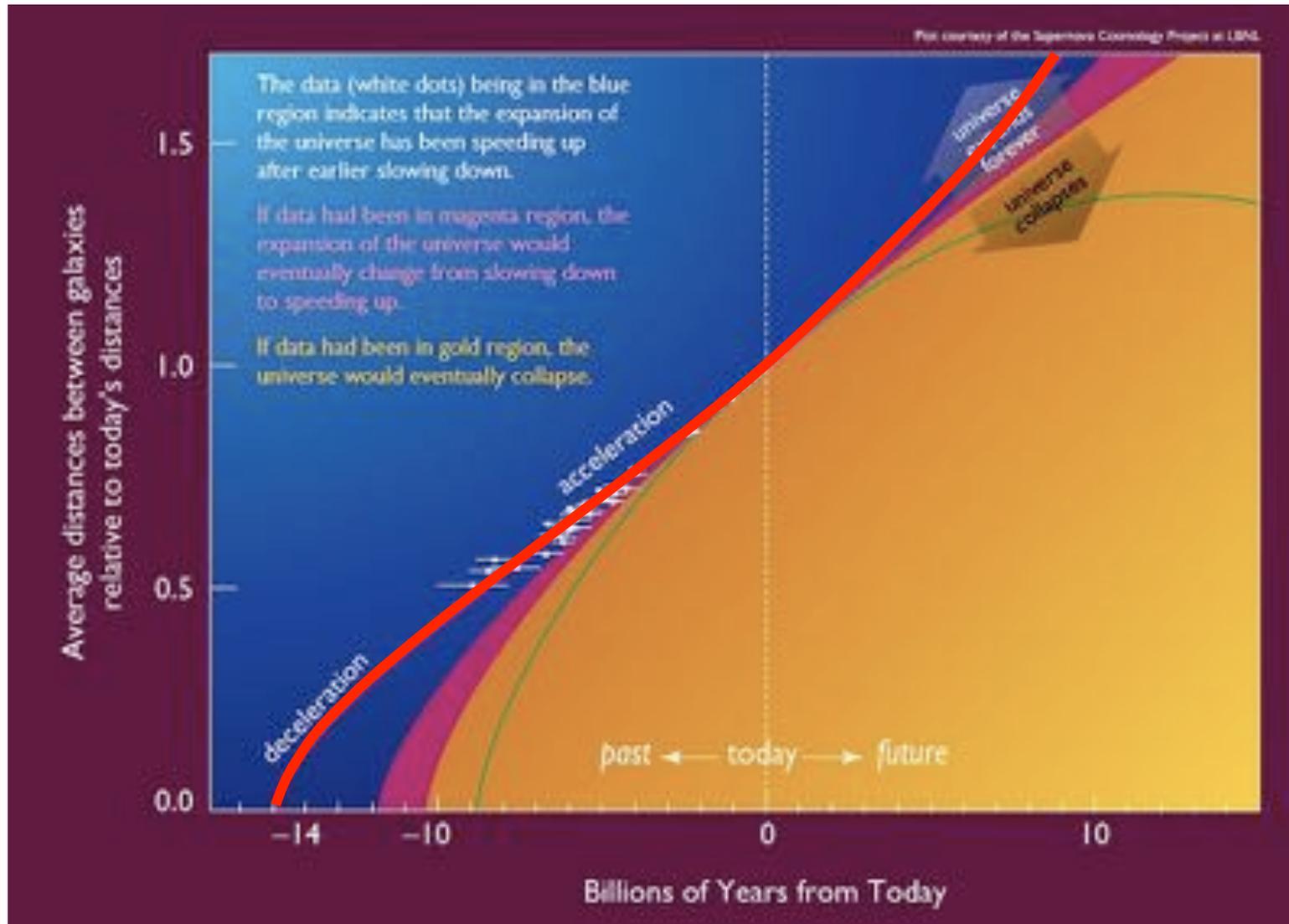
Supernovae de type Ia

le message des chandelles standart SNIa

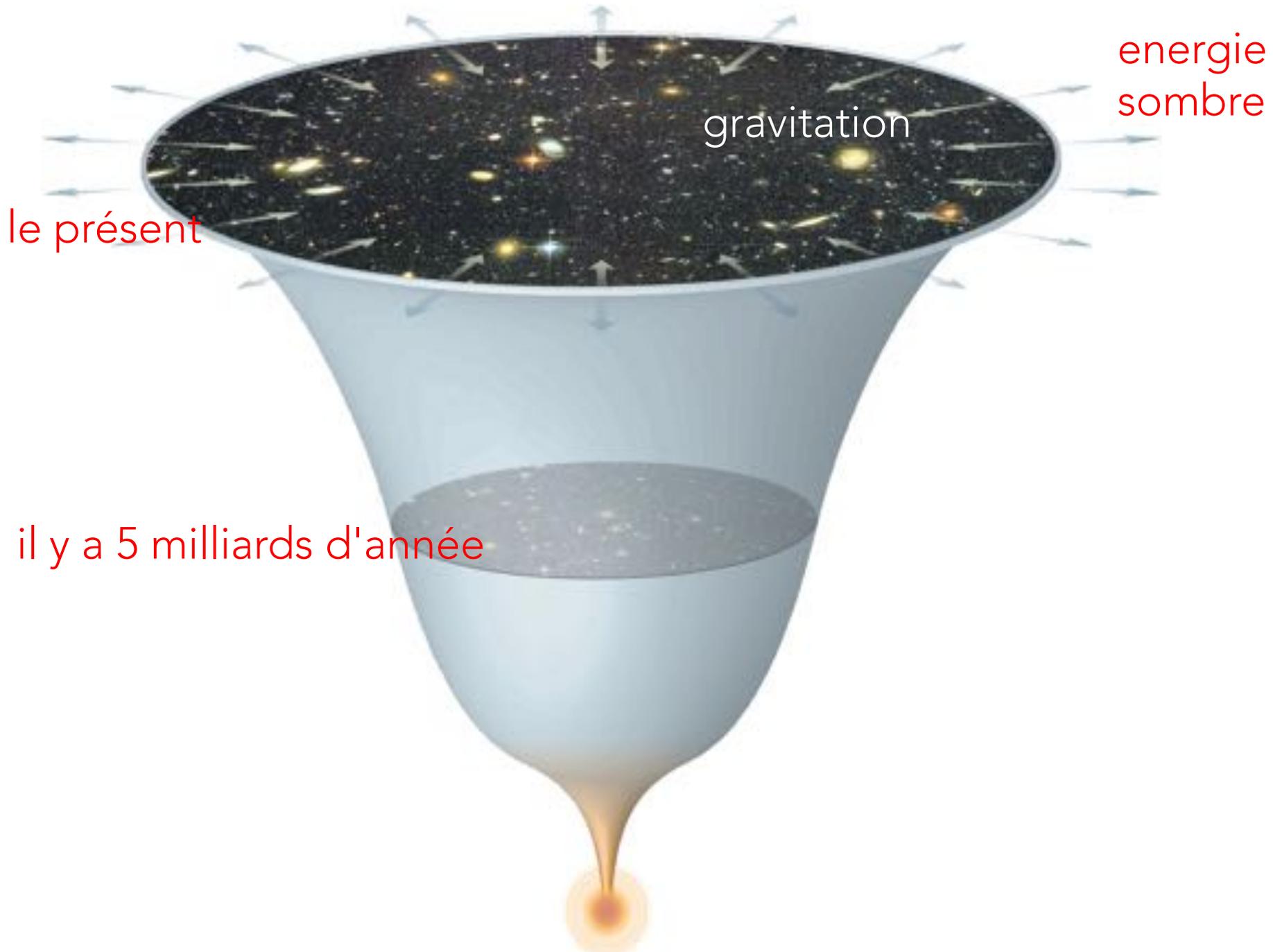


From left, Adam Riess, Saul Perlmutter and Brian Schmidt shared the Nobel Prize in physics awarded Tuesday.

L'histoire de l'expansion universelle



Expansion accélérée -> nouvelle forme d'énergie : l'énergie noire



le présent

gravitation

energie
sombre

il y a 5 milliards d'année

Energie noire

- 70% du bilan énergétique
- répartie de façon extrêmement uniforme dans l'univers
- propriétés opposées à celles de la matière ordinaire (répulsion gravitationnelle)
- ressemble beaucoup à la constante cosmologique d'Einstein

Matière noire

- 25% du bilan énergétique
- répartie à peu près comme les étoiles, mais en plus diffus
- propriétés analogues à la matière ordinaire, quoique elle interagisse moins
- longue liste de candidats rejetés
- liste encore plus longue de candidats imaginés

dans la salle de contrôle de l'Univers

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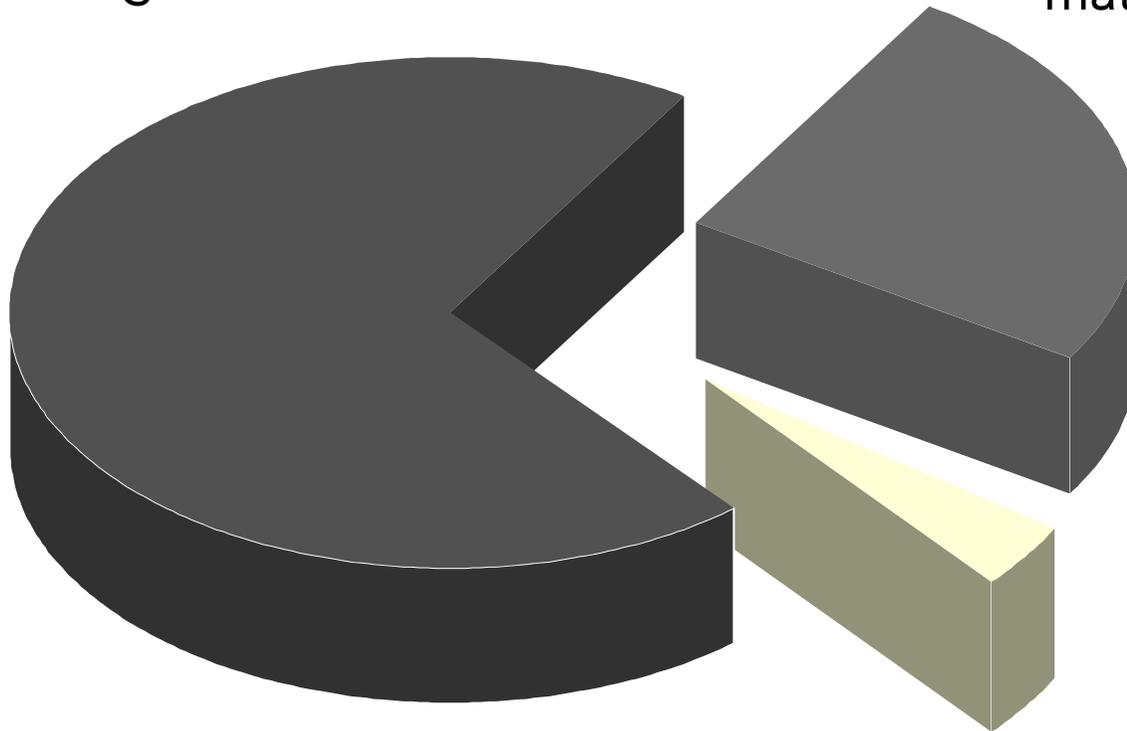
Λ

la constante cosmologique

l'Univers est constitué de :

énergie sombre 68.3 %

matière noire 26.8 %



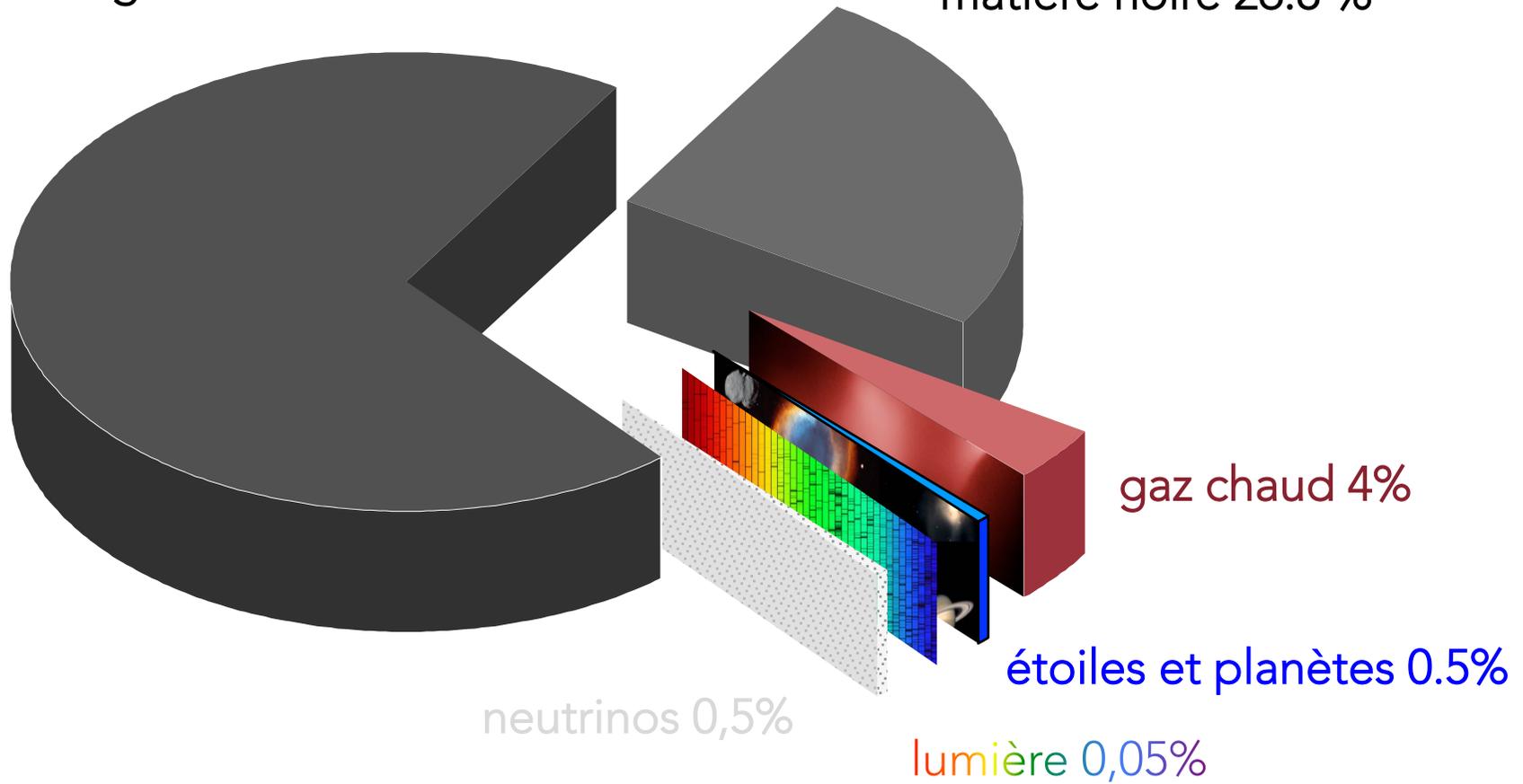
matière baryonique 4.9 %

selon les données Planck/ESA (3/2013)

l'Univers est constitué de :

énergie sombre 68.3 %

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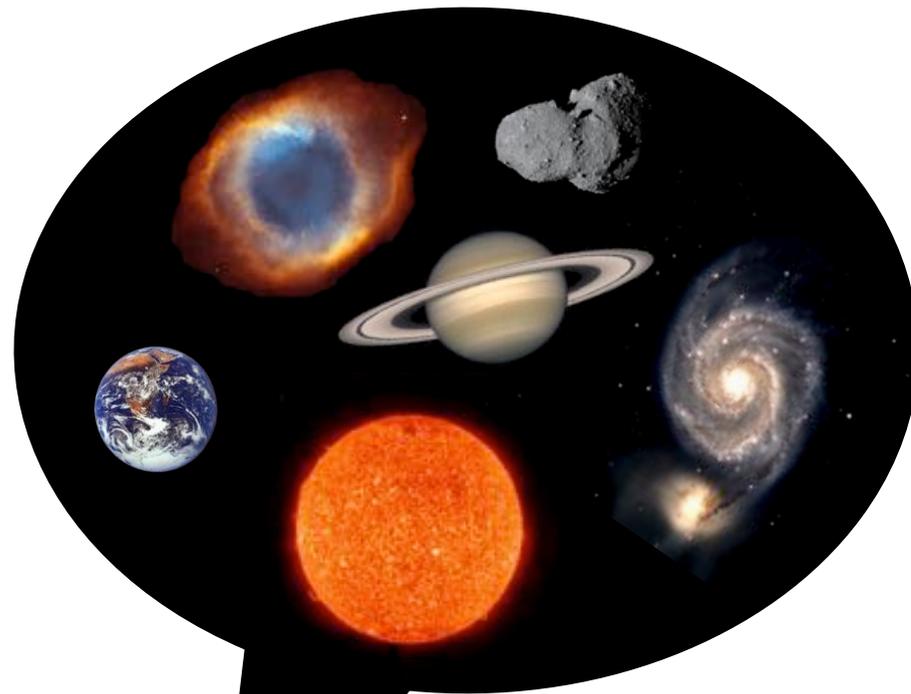
la pointe de l'iceberg :

étoiles et planètes

gaz chaud

lumière

neutrinos



la pointe de l'iceberg

matière 4.9 %

matière noire 26.8 %

énergie sombre 68.3 %

